## Laboratory of Theoretical Physics

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THE SIMULATING OF MULTIPLE PRODUCTION PROCESS he $\exists$ Top, $1958,735, b 6(12)$, c $1426-1434$.

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## ABSTRACT

An idea of constructing the "table of random stars" capable of reproducing the theoretical concepts about the multiple production of elementary particles in the form suitable for comparing with experiments is discussed. The possibility of making such a table in the energy range up to 10 BeV is shown.

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A majority of papers on rae multiple production theory is devoted to the solution of two problems: to the calculation of the statistical weights of different reactions and to the obtaining of momentum distributions. Apart from this there remains a problem about the angular distributions and correlations of various kinds between the directions and velocities of the secondary particles. Meanwhile, its solution would make it possible to judge more definitely about the applicability of the theory to the explanation of the multiple production process. Indeed, the correlations between the directions of the particle motion must depend upon the character of the interaction between particles at the moment of their production ; in the first Fermi model of the statistically independent secondary particles the correlations must depend only upon the limitations due to the conservation laws; in the concept of the resonance interaction of a nucleon and meson accepted now the magnitude of the correlations must be greater. Comparison of the quantitative expressions for the correlations with experiment would make, perhaps, the character of the resonance interaction more exact or would outline the limits of the isobar theory.

The qualitative evaluation of the correlations "forwardbackward" was made by Fermi/I/ taking into account the law of angular momentum conservation. However, the complexity of the necessary calculations is the obstacle for the quantitative conclusions. If they could be performed under simple assumptions then the complication of the form of the interaction matrix /the natural way of developing the theory/ would make the calculation more difficult. Under these conditions the simulating of the multiple production process would be an evident way out. However, since our idea about the multiple proauction mechanism is too vague to find a large scale process with similar features it is necessary to make use of the numerical simulating. Such a simulating must be capable of reproducing the whole class of the assumptions about the form of the interaction.

An idea of such simulating was suggested by M.I. Podgoretsky and M.Danysz in the form of a "table of random stars" according to an isobar model. This idea consists in the tabulating of the random numbers which obey the same regularities as the momenta of the secondary particles in the isobar theory. The table of random stars must contain a number of lines ; each of them is the description of one event of particle production; i.e. it contains the magnitudes
and directions of the momenta of secondary particles. The statistical treatment of such a table will yield the same data about the multiple production process-the statistical weights, momentum, angular and charge distribution etc, as well as the treatment of the real stars in the photoemulsion or in the chamber. The comparison of such data will answer the question about the agreement of the theory with experiment.

The table of randon stars includes in principle all the data which can be obtained from the real stars. This is the advantage over other calculation methods. Any distributions or correlations which can be obtained from the stars in the photomulsion may be repeated in the corresponding table. The shortcoming of the method is a limited accuracy, a great deal of calculations necessary for making up and analyzing the table and the impossibility of obtaining the analytic expressions for the results.

It is shown in this paper how to realize the above idea by means of the computer. When there are $5-6$ secondary particles and under some simplifying assumptions one may draw a table of IOO-200 lines using the abovementioned method by means of the hand-operated calculation.

The method under consideration takes into account the energy and momentum conservation in the reactions. It can be generalized over a rather wide class of ideas about the particle interaction in their multiple production. This enables to give a positive answer to the question about the possibility of the multiple production simulating.

Paragraph I concerns drawing up the table of the reactions in which the number of secondary particles is equal to three, the Fermi model being oonsidered correct. This simple example illustrates the main features of the simulating method under consideration.

In Paragraph 2 the method is generalized for $n$ secondary particles and for the arbitrary form of the square $F$ of the matrix interaction element; some details which make the calculations simpler are given in brief. At the same time some kinematic characteristics of the system of $n$ particles are derived.

In Parggraph 3 the procedure for drawing up the table of random stars suitable for the comparison with experiment is outiined.

## § I. TABLE OF ThREE PARTICLE REACTIONS

Consider the reaction of two particle interaction with the total energy $E$ and zero momentum. As a result of this reaction three particles with masses $m_{i}$, momenta $\vec{P}_{i}$ and total energies $e_{i} / i=I, 2,3 /$ are produced. According to Fermi theory the probability of this reaction is proportional to

$$
W(E)=\int d^{3} \overrightarrow{p_{1}} d^{3} \overrightarrow{p_{2}} d^{3} \overrightarrow{p_{3}} \delta\left(\sum_{1}^{3} e_{i}-E\right) \delta\left(\sum_{1}^{3} \overrightarrow{p_{i}}\right)
$$

The faotor $\delta\left(\sum_{i}^{3} e_{l}-E\right) \delta\left(\sum_{1}^{3} \vec{p}_{i}\right)$ is the momentum distribution density in the momentum nine-dimensional space; tne problem of simulating consists in realizing such distribution. But the narrow stripes along the intersection of the surfaces $\sum_{1}^{3} e_{i}-E=0 \quad$ and $\quad \sum_{i}^{3} \vec{p}_{i}=0 \quad$ in the nine-dimen-
sional rectangle have too small volume for the points with the coordinates $\vec{P}_{i}$ chosen at random to hit these stripes sufficiently often. But the integral obtained from/I.I/ by simple calculations

$$
W(E) \int_{D_{3}} d p_{1} \alpha_{p_{2}} \cdot 8 \pi^{2} p_{1} p_{2}\left(E-e_{1}-e_{2}\right)
$$

makes it possible to realize the necessary momentum distribution. The integration region $\mathscr{D}_{3}$ in 1t/Fig. I/ is bounded by the curves $\cos \theta= \pm 1 / Q$ is the angle between $\vec{p}_{1}$ and $\vec{P}_{2}$
/. Their equations are

$$
\varepsilon^{+}=0, \varepsilon^{-}=0
$$

where

$$
\varepsilon^{ \pm}=e_{1}+e_{2}+\sqrt{\left(\rho_{1} \pm p_{2}\right)^{2}+m_{3}^{2}}-E
$$

It follows from /I. $2 /$ that the density distribution of the probability to have the momentum $p_{1}$ in $d p_{1}$ and at the same time $p_{2}$ in $d p_{2} / 1 . e$. the density of points $M\left(p_{I}\right.$, $p_{2}$ ) in $D_{3} / 1 s$ proportional to the function

$$
W\left(p_{1}, p_{2}\right)=p_{1} p_{2}\left(E-e_{1}-e_{2}\right) .
$$

The points $M$, according to /I.5/, are distributed in $D_{3}$ not uniformly. We shall have to be concerned with the nonuniform distributions below. Let us consider the ways of their realization with the heip of the uniform ones.

To obtain the random distributed numbers $\alpha$ in / a, b/with the density $f(\alpha) / F 1 g$. 2/ there are two possibilities: $A /$ rejection technique and $B / a \operatorname{direct"}$ method $/ 5 /$.
A. Two numbers are picked simultaneously : $\alpha$ uniformis in $/ a, b /$ and $\beta$ uniformiy in / $0, M^{\prime} /$, where the constant $M^{\prime} \geqslant f(\alpha)$ for all $\alpha$ on ( $a, b$ ). Those pairs $\alpha, \beta$ for which $f(\alpha)<\beta$, are rejected, then the remaining values of $\alpha \quad$ wili be distributed with the density $\mathcal{X}(\alpha)$. Indeed, at great number of samples the number of values of $\alpha$ close to $\alpha=\alpha_{1}$ tends to be equal to the number of $\beta<f\left(\alpha_{1}\right)$, i.e. to the quantity $f\left(\alpha_{1}\right) / M^{\prime}$ since $\beta$ are distributed over the ordinate $\alpha=\alpha_{1}$ uniformly.

This method is also suttable for many-dimensional distributions, however, at sharp maxima of the function f(ل) number of the discarded pairs $\alpha, \beta$ may be found too great; if $f(\alpha)$ tends to infinity then this method is not suitable at all/ or gives only approximate results/.
B. $\quad \beta$ is selected uniformly in $/ 0, I /$ and the equation $\int_{0}^{\alpha} f(\alpha) \alpha \alpha=\beta$ is solved each time. Then the distribution density of the solutions of this equation is proportional to.
$f(\alpha)$ since $\quad \quad \quad d n / d \alpha=\left(d_{n} / d \beta\right) \cdot f(\alpha) \sim f(\alpha)$
Through this method is applicable only to the one- dimensional distributions, it is suitable for infinite densities and is not connected with the necessity of rejecting the ohosen values of $\beta$.

The two-dimensional distribution $/ I .5 /$ is convenient to realize using the rejection technique. For this it is necessary to know the maximum of $W\left(P_{1}, P_{2}\right)$ for $D_{3}$. one oan easily find that the values $P_{1,2}=\bar{P}_{1,2}$, at whioh this maximum is achieved are the roots of
the system

$$
p_{1}^{2} / e_{1}=p_{2}^{2} / e_{2}=E-e_{1}-e_{2}
$$

If they are found inside $D_{3}$. Otherwise $\bar{P}_{1,2}$ must lie on the boundary $D_{3}$, on the curve $\varepsilon^{-}=0$, and satisfy the equation system

$$
\frac{P_{2}\left[E-e_{1}-e_{2}-\left(P_{1}^{2} / e_{1}\right)\right]}{P_{1}\left[E-e_{1}-e_{2}-\left(P_{2}^{2} / e_{2}\right)\right]}=\frac{P_{2}-\left(E-e_{2}\right)\left(P_{1} / e_{1}\right)}{P_{1}-\left(E-e_{1}\right)\left(P_{2} / e_{2}\right)}, \quad \varepsilon^{-}=0 . \quad \text { /I. 7-I.8/ }
$$

At high energies $E$ one has to solve /I.6/, at low energies
-/I.7-I.8/.
A pair of values $P_{I}, P_{2}$ determines the momenta of all three particles by their magnitudes and directions. Therefore, the sequence of obtaining one line of three -particle star table corresponding to the Fermi model 19 the following :

I/ The random numbers $p_{I}, p_{2}$ in the intervals ( $\left.O, \widetilde{p}_{1} \max \right),\left(0, \widetilde{P}_{2} \max \right)$ are chosen uniformiy: Here

$$
\tilde{p}_{1 \max }=\left\{\left[E^{2}-\left(m_{1}+m_{2}+m_{3}\right)^{7}\right]\left[E^{2}-\left(m_{1}-m_{2}-m_{3}\right)^{2}\right]\right\}^{1 / 2} / 2 E
$$

and an analogous form holds for $\tilde{P}_{2} \max$
2/ it would be verified whether the point ( $p_{I}, p_{2}$ ) lies in $D_{3}$. One may do this in the simplest way by checking the fulfillment of the inequalities

$$
\varepsilon^{-}<0, \varepsilon^{+}>0 \quad / I \cdot I 0 /
$$

If they are satisfied then one oan always find suoh an angle $\theta$ between $\vec{p}_{1}$ and $\vec{p}_{2}$ that

$$
e_{1}+e_{2}+\sqrt{\left(\overrightarrow{p_{1}}+\overrightarrow{p_{2}}\right)^{2}+m_{3}^{2}}=E
$$

If /I.II/ is not satisfied then the pair ( $p_{I}^{\prime}, p_{2}$ ) is rejected, instead of it a new one is selected.

3 in $\left(O, W_{\text {max }}=W\left(\bar{P}_{1}, \bar{P}_{2}\right)\right.$ ) the random number $\beta$ is pioked and the inequality

$$
\beta \leqslant w\left(p_{1}, p_{2}\right)
$$

is verified.
If it is not fulfilled then the pair / $p_{I}, p_{2} /$ is discarded, if it is fultilled then : 4/ from /I.II/

$$
\cos \theta=\left(p_{3}^{2}-p_{1}^{2}-p_{2}^{2}\right) / Q p_{1} p_{2}
$$

is caloulated.
The momenta of three particles may now be considered equal to

$$
\begin{gathered}
\overrightarrow{p_{1}}=\left\{p_{1} \cos \psi,-p_{1} \sin \psi, 0\right\}, \overrightarrow{p_{2}}=\left\{p_{2} \cos (v-\psi), p_{2} \sin (v-\psi), 0\right\} \\
\overrightarrow{p_{3}}=-\overrightarrow{p_{1}}-\overrightarrow{p_{2}},
\end{gathered}
$$

where $\psi$ is an random angle unitormp distributed in $/ 0,2 \pi / \%$ It is introduoed in order the directions of any momentum to be isotropio in the plane ( $\vec{p}_{1}, \vec{p}_{2}, \vec{P}_{3}$ ). .

The table each line of which is oomputed by /I.I4/ will be the table of plane stars. When making the analysis of the experiment this will require to bring all three- particle
stars to one plane and only after this to compare the statistics of the experiment with that of the Table. To make the Table of three- particle stars which do not lie in the same plane It is, evidently, necessary to rotate the end of the vector of the unit normal to the plane $/ \overrightarrow{p_{1}}, \overrightarrow{p_{2}}, \overrightarrow{p_{3}} /$ randomly about a certain axis.
The uniform distribution is most simply achieved if the cosine $\eta$ of the angle between the normal and the axis $0 Z$ is chosen uniformily in ( $-I, I$ ) as well as the angle $\varphi$. between the axis. $O X$ and the projection of the normal to the plane XOY uniformly in ( $0,2 \boldsymbol{J}$ ).

Then the number of hits of the normal to the element of the surface / $P, \varphi / w i l l$ appear proportional to the area of this element $\left(\alpha_{\eta} \alpha y\right)$.

The angles $\varphi$ and $\cos ^{-1} \gamma$ together with the angle $\psi$ form three Eiler angles describing the transformation of the system $O X Y Z$ in $O x^{\prime} y^{\prime} Z^{\prime / 2 /}$. In Table of solid stars the components of the momenta must therefore be calculated by the formula $\vec{P}^{\prime}=A \vec{P}$, , where $A$ is a well-known transformation matrix

$$
A=\left(\begin{array}{ccc}
\cos \varphi \cos \psi-\eta \sin \varphi \sin \psi & \sin \varphi \cos \psi+\eta \cos \varphi \sin \psi & \sqrt{1-\eta^{2}} \sin \psi \\
-\cos \varphi \sin \psi-\eta \sin \varphi \cos \psi & -\sin \varphi \sin \psi+\eta \cos \varphi \cos \psi & \sqrt{1-\eta^{2}} \cos \psi \\
\sqrt{1-\eta^{2}} \sin \varphi & -\sqrt{1-\eta^{2}} \cos \varphi & \eta
\end{array}\right) / I \cdot 15 /
$$

While the momenta may be put equal to

$$
\vec{p}_{1}=\left\{p_{1}, 0,0\right\}, \quad \overrightarrow{p_{2}}=\left\{p_{2} \cos \theta, p_{2} \sin \sigma_{0} 0\right\}
$$

Another means for transforming the plane stars into those uniformiy distributed in space is stated in § 2. The estimate of the efficiency of the supposed methods for obtaining the stars is given there too.

## 2. TABLE OF REACTIONS WITH $n$ SECONDARY PARTICLES

General case. Consider the case of the appearance of $n$ secondary partioles in the interaction which is characterized by the square of the matrix element $\vec{F}\left(\vec{P}_{1}, \ldots, \vec{P}_{n}\right)$ tending to infinity nowhere. $F$ may also depend upon the onergy $E$ and the parameters of the primary particles. But it will be not essential further. The momentum part of the expression for the statistical weight has the form

$$
W(E, \vec{P})=\int d^{3} \vec{p}_{1} \ldots d^{3} \vec{p}_{n} F\left(\vec{p}_{1}, \ldots, p_{n}\right) \delta\left(\sum_{1}^{n} e_{i}-E\right) \delta^{\prime}\left(\sum_{1}^{n} \vec{p}_{i}-P\right),
$$

where $\vec{P}$ is the total momentum of the system. The integration region over the first /x-I/ momenta $\vec{P}_{1}, \ldots, \vec{P}_{x-1}$ will be denoted by Dr, at the same time no restriations except the conservation laws are imposed upon the rest momenta $\vec{p}_{k}, \ldots, \vec{P}_{r}$. Let the integration region over $\vec{p}_{k}$ at fixed $\overrightarrow{P_{1}}, \ldots, \overrightarrow{P_{k-1}}$ and arbitrary $\vec{p}_{k+1}, \cdots, \overrightarrow{P_{n}} /$ with account of the conserv-
ation laws/ be $d \kappa$. Introduce special notations for energy and momentum which are left for the particles $\kappa, \ldots, n$ at the fixed momenta of the particles $1, \ldots, k-1$

$$
\begin{align*}
& E_{k}=E-\sum_{1}^{k-1} e_{i} \\
& P_{x}=\sum_{1}^{\alpha-1} \vec{p}_{i}-\vec{P}
\end{align*}
$$

as well as for characteristic /"limiting"/ energy and momentum of the particle $k$ in the centre of mass system of particles $k, \ldots, n$ at fixed $\vec{p}_{1}, \ldots, \vec{p}_{k-1}$ :

$$
\begin{aligned}
& E_{\kappa}^{*}=\left(M_{\kappa}^{2}+m_{\kappa}^{2}-M_{x}^{2}\right) / 2 M_{\kappa} \quad(\kappa=1, \ldots, n-1), \\
& E_{n}^{*}=\left(M_{n-1}^{2}-m_{n-1}^{2}+m_{n}^{2}\right) / 2 M_{n-1}, \\
& P_{\kappa}^{*}=\sqrt{E_{\kappa}^{* 2}-m_{\kappa}^{2}}=\sqrt{\left(M_{\kappa}^{2}-m_{\kappa}^{2}-M_{\kappa}^{2}\right)^{2}-\left(2 M_{\kappa} M_{\kappa}\right)^{2}} / 2 M_{\kappa},
\end{aligned}
$$

where $M_{K}$ is the effective mass of a system of particles $\kappa, \ldots n$, whereas $M_{k}$ is the mass of the compound particle composed from $k+1, \ldots, n$ :

$$
\begin{align*}
& M_{\kappa}^{2}=E_{\kappa}^{2}-P_{\kappa}^{2} \\
& M_{\kappa}=\sum_{\kappa+1}^{n} m_{i}
\end{align*}
$$

We shall set the components of the momentum $\overrightarrow{P_{k}}$ in the spherical coordinate system $\vec{p}_{K}=\left\{p_{K}, \vartheta_{K}, \varphi_{K}\right\}$, connected with the sum of the momenta $\vec{p}_{I}, \ldots, \vec{P}_{k-1} \quad$ already chosen. Let the polar axis of the system ta be directed along $\overrightarrow{P_{k}} ; \vartheta_{n}$ is the angle between the vectors $\overrightarrow{P_{k}}$ and $\overrightarrow{P_{K}}$, the azimuth $\varphi_{K}$ is counted off the vertical plane in which $\overrightarrow{P_{K}}$ lies /Fig.3/. This coordinate system is convenient because the module of the sum of momenta $\overrightarrow{\boldsymbol{P}_{k}}$ and $\overrightarrow{p_{k}}$ does not depend upon the angle $\varphi_{\kappa}$

$$
p_{k+1}^{2}=p_{k}^{2}+p_{k}^{2}+2 p_{k} p_{k} \cos \vartheta_{k}
$$

For recounting the components of the momentum into the rectangular coordinate system one may use the formulas following from Fig. 3.

$$
\begin{align*}
& \vec{P}_{k}=\left\{x_{\kappa}, y_{\kappa}, Z_{\kappa}\right\} \\
& R_{k}=\left(x_{\kappa}^{2}+y_{k}^{2}\right)^{1 / 2} \text {, } \\
& \left.x_{k}=p_{k} \left\lvert\, \frac{x_{k}}{P_{k}} \cos \theta_{k}+\left(\frac{Z_{k}}{P_{k}} \frac{x_{k}}{R_{k}} \cos \varphi_{K}-\frac{y_{k}}{R_{k}} \sin \varphi_{k}\right) \sin \theta_{K}\right.\right], \\
& y_{k}=P_{k}\left[\frac{y_{k}}{P_{k}} \cos \theta_{k}+\left(\frac{Z_{k}}{P_{k}} \frac{y_{k}}{R_{k}} \cos \varphi_{k}+\frac{x_{k}}{R_{k}} \sin \varphi_{k}\right) \sin \theta_{k}\right] \text {, } \\
& Z_{k}=P_{k}\left[\frac{Z_{k}}{\rho_{k}} \cos \theta_{k}-\frac{R_{k}}{\rho_{k}} \cos \varphi_{k} \sin \theta_{k}\right] .
\end{align*}
$$

Now let us transform /2.I/. After the integration over $\vec{\rho}_{n}$ and over $v_{n-1}$ ס-functions in the integrand vanish:

$$
W(E, \vec{P})=\int d^{3} \vec{p}_{1} \ldots d^{3} \vec{p}_{n-2} d p_{n-1} d \varphi_{n-1} p_{n-1} \frac{E_{n}}{P_{n-1}} \mathcal{F}\left(\vec{p}_{1}, \ldots, \vec{p}_{n-1},-\vec{p}_{n-1}-\vec{p}_{n-1}\right), \quad 12.131
$$

Where the vector $\overrightarrow{\rho_{n-1}}$ is of the form

$$
\vec{p}_{n-1}=\left\{p_{n-1} \cos ^{-1} \frac{E_{n}^{2}-m_{n}^{2}-p_{n-1}^{2}-p_{n-1}^{2}}{2 P_{n-1} p_{n-1}}, \varphi_{n-1}\right\},
$$

Whereas the integration region $D_{n}$ is limited by the surfaces $\cos v_{n-1}= \pm$, i.e.,

$$
\sqrt{\left(P_{n-1} \pm p_{n-1}\right)^{2}+m_{n}^{2}}=E_{n-1}-e_{n-1} .
$$

In principle expression /2.I3/ makes it possible to solve the problem of the realization of the distribution required by the model with the interaction $F$. In $/ 2 . I 3 /$ there are no dependent integration variables, therefore the sets of values of $\vec{P}_{1}, \ldots, \vec{P}_{n-2}, P_{n-1}, \varphi_{n-1}$ must satisfy certain inequalities instead of equalities/ as distribution / $2.1 /$ would have required/, what increases the efficiency of sampling. However, the amount of the calculations necessary for creating the distribution over $\overrightarrow{P_{1}}, \ldots, \overrightarrow{P_{n-2}}, P_{n-1}, Y_{n-1}$ of the density $p_{n-1} E_{n} F / \rho_{n-1}$ is still rather great. All further transformations must be aimed at reducing this amount of calculations. An evident way for this is to perform in $/ 2.13 /$ the maximum possible number of integrations. Due to this the dimension of the integration region falls increasing thereby the ratio of the volume of the region to the volume of the rectangle circumscribed. However, this aim cannot be achieved if the form of the function $F$ is arbitrary. Having in view to suggest the method applicable for any F we shall choose another way of reducing the calculations, and namely, we present $d^{3} \overrightarrow{p_{k}}$ as $p_{\kappa}^{2} d p_{\kappa} d \cos v_{k} d \varphi_{\kappa}$ using the spherical ooordinate system described above. Then the distribution density over $p_{1}, \cos \theta_{1}, \varphi_{1} ; p_{2}, \ldots ; \ldots ; \ldots, \varphi_{n-2} ; p_{n-1}, \varphi_{n-1}$ is of the form:

$$
\phi\left(\rho_{1}, \cdots, \varphi_{n-1}\right)=\rho_{n-1} \frac{E_{n}}{P_{n-1}} \mathcal{F} \cdot \prod_{k=1}^{n-2} \rho_{\kappa}^{2}
$$

where the arguments $F$ are the same as in $/ 2.3 /$. It is essential that the regions $\mathcal{D}_{2}$, $\ldots, \mathcal{D}_{\kappa}, \ldots, D_{n}$ of the allowed values of $\overrightarrow{p_{1}} ; \ldots, \overrightarrow{p_{1}}, \ldots, \overrightarrow{p_{k-1}} ; \ldots ; \overrightarrow{p_{1}}, \ldots, \overrightarrow{p_{n-1}}$ are independent of $Y_{1}, \ldots, Y_{n-1}$. Let us prove this.

Let the momenta $\overrightarrow{P_{1}}, \cdots, \vec{P}_{k-1}(K \leqslant n-2)$ be fixed. Find the region $d_{k}$ of the change of the momentum $\overrightarrow{P_{k}}$. For the remaining particles $k, k+1, \ldots, n$ the equations

$$
\begin{align*}
& \sum_{k}^{n} e_{i}=E_{k} \\
& \sum_{k}^{n} \overrightarrow{p_{i}}=-\overrightarrow{p_{k}}
\end{align*}
$$

are correct.

It is known $/ 3 /$ that the limiting values of the momentum of one of the particles at their multiple production are obtained when all other secondary particles move as a whole, i.e., as a particle of mass equal to their total mass. For particle $k$ this mass is equal to $\mu_{k}$ /see /27//. The conservation laws for two particles mith masses $m_{k}$ and $\mu_{k}$, with the total energy $E_{K}$ and the momentum $-\overrightarrow{P_{K}}$ are written as follows

$$
\sqrt{p_{k}^{2}+m_{k}^{2}}+\sqrt{\left(-\vec{P}_{k}-\vec{P}_{k}\right)^{2}+\mu_{k}^{2}}=E_{k}
$$

This is an equation of a surface limiting the region of possible positions of the vector end
$\overrightarrow{p_{k}}$. The position and form of the region $\mathcal{C}_{\kappa}$ mas be most explicitly obtained from the graphic representation of the conservation laws /4/. The sphere $P_{k}^{*}=$ Const. in the center of mass system is deformed when passing into the laboratory system into the ellipsoid of revolution extended along the direction of mutual motion of two frames of references. In the considered case the center of the ellipsoid is the point $o\left(-E_{k}^{*} \vec{P}_{k} / M_{k}\right)$, the major axis has the length $2 p_{K}^{*} E_{K} / M_{K}$. It is directed along the vector $\vec{p}_{K}$ and $1 s$ an axis of revolution, the semi-minor axis is equal to $P_{k}^{*} / F 1 g .4 /$. Here $E_{k}^{*}$ and $P_{k}^{*}$ are defined by /2.4/ -/2.5/.

Now the region $\mathscr{D}_{K}$ may be found. Sinoe the energy $E_{k}^{*}$ of the particle $k$ in the center of mass system of partioles $k, k+I, \ldots, n$ is not less than the mass of this particle. $\mathrm{m}_{\mathrm{k}}$ then /see/2.4//

$$
M_{k} \geqslant \mu_{k-1}(k \leq n-1)
$$

: Condition $/ 2.18 /$ is not only a necessary restriction upon $\vec{P}_{1}, \ldots, \overrightarrow{P_{k-1}}$, but a sufficient one: when it is fulfilled there may be always found such $\vec{P}_{k}, \ldots, \vec{P}_{n}$ that the conservation laws will be satisfied. It is possible for instance to direct the partioles. $k+I, \ldots, n$ all together (in the frame of reference where $\vec{P}_{\kappa}=0$ ) in one direction whereas the particle $k$ in the opposite one.

But. $M_{K}$ depends only, upon the modulus $/ 2.6 /$ which in its turn 18 independent of $\varphi_{K}$ 12.8/. Therefore the form of the region $\mathscr{D}_{K} / 2.18 /$ is independent of $\varphi_{K}(\kappa \leq n-1)$. For $D_{0}$ it follows from /2.15/.

The independence of the outlines of the regions $D_{K}$ of $Y_{k}$ for all $k$ from 1 up to $n$ makes it possible to reduce the dimension of each $D_{k}$ as much as one and a half times. So, the dimension $D_{n} w 111$ be $2(n-I)$ instead of $3(n-I)$. Thereby the ratio of the volume of the region to the volume of circumscribed rectangle of the same dimension considerably increases.

Write down in more detail the limits restricting the dimensions of $D_{k}$ and of the rectangle circumscribed. It is clear from Fig. 4 that at fixed $\overrightarrow{P_{1}}, \ldots, \overrightarrow{P_{K-1}}$ the limits of the change of $p_{k}$ are given by the formulas

$$
\begin{aligned}
& P_{\text {max }}^{(\alpha)}=\left(E_{k}^{*} P_{k}+P_{k}^{*} E_{k}\right) / M_{k} \quad \text { at } k=1, \ldots, n-1, \\
& P_{\text {min }}^{(k)}=\left\{\begin{array}{cc}
\left(E_{k}^{*} P_{k}-P_{k}^{*} E_{k}\right) / M_{k} & \text { at } E_{k} / M_{k} \geqslant E_{k}^{*} / m_{k} \text { and } k<n-1 \\
0 & \text { at } E_{k} / M_{k} \leq E_{k}^{*} / m_{k} \text { and } k<n-1 / 2.19 /
\end{array}\right. \\
& P_{\min }^{(n-1)}=\left|E_{n-1}^{*} P_{n-1}-P_{n-1}^{*} E_{n-1}\right| / M_{n-1}
\end{aligned}
$$

/The restriotions upon $P_{n-1}$ differ from those upon $P_{k} / \kappa<n-1 /$ because the system of particles $n-I$ and $n$ is kinematically definite/.

The limits of the change of $\cos \theta_{k}$ at the fixed $\vec{P}_{,}, \ldots, \vec{P}_{k-1}, P_{k}$ have the following form

$$
\begin{aligned}
& \left(M_{\kappa} E_{k}^{*}-E_{\kappa} e_{\kappa}\right) / P_{k} p_{k} \geqslant \cos \theta_{k} \geqslant-1 \text { at } E_{\kappa} / M_{\kappa} \geqslant E_{\kappa}^{*} / m_{k} \text { or } \\
& \text { at } E_{\kappa} / M_{\kappa} \leqslant E_{\kappa}^{*} / m_{\kappa} \text { and } p_{\kappa} \geqslant\left(p_{k}^{*} E_{\kappa}-E_{\kappa}^{*} \rho_{\kappa}\right) / M_{k} \text {, } \\
& \text { /2.20/ } \\
& 1 \geqslant \cos \vartheta_{k} \geqslant-1 \text { at } E_{k} / M_{K} \leqslant E_{k}^{*} / m_{k} \text { and } \rho_{k} \leqslant\left(p_{k}^{*} E_{K}-E_{\kappa}^{*} P_{k}\right) / M_{k} \text {. }
\end{aligned}
$$

The boundaries of the circumscribed rectangle are obtained by the same oonsiderations but under the assumption that no momentum is fixed. For the momentum of the particle $k$ ( $k=I, \ldots, n$ ) a following chain of formulae is obtained

$$
\begin{gather*}
M^{2}=E^{2}-P^{2}, \\
e_{k}^{*}=\left[M^{2}+m_{k}^{2}-\left(\sum_{i \neq k}^{n} m_{i}\right)^{2}\right] / 2 M \\
P_{k}^{*}=\sqrt{e_{k}^{* 2}-m_{k}^{2}} \\
\tilde{P}_{k \text { max }}=\left(E P_{k}^{*}+\rho e_{k}^{*}\right) / M \\
\tilde{P}_{k \min }=\left\{\begin{array}{cl}
\left(P_{e_{k}^{*}}-E p_{k}^{*}\right) / M & \text { at } E / M \geqslant e_{k}^{*} / m_{k} \\
0 & \text { at } E / M \leq e_{k}^{*} / m_{k}
\end{array}\right.
\end{gather*}
$$

The limits of the ohange of $\cos \theta_{\pi}$ are $\pm I$ /except $\cos \theta_{1 /}$.
So, it is possible to outline the following order of sampling $n$ - partiole reaotions characterized by the square of the matrix element $F$.
$p_{I}$ and cos $\vartheta_{1}$ are seleoted in $\left(\tilde{P}_{\text {min }}, \tilde{P_{1} \max }\right)$ and $(-I,+I)$, then $p_{2}$ and oos $\theta_{2}$ are chosen in $\left(\tilde{P}_{2 \min }, \tilde{P}_{2 \text { max }}\right)$ and $(-I,+I)$, and the fulfillment of $(2.18)$ for $k=3$ is
verified etc... $p_{k}$ and $\cos \varphi_{r}$ are chosen $\ln \left({\overline{P_{k}}}_{\min }, \bar{P}_{\kappa \text { maxe }}\right)$ and $(-I,+I), E_{k+I}$ and $P_{k+I}$ are calculated by (2.2) and (2.8), and (2.I8) is verified; when (2.I8) is not fulfilled the sampling begins again from $p_{I}$, otherwise one passes in to ( $k+I$ ) etc. up to $k=n-I$, when only $P_{n-1}$ are picked in ( $P_{\text {min, }}^{(n-1)} P_{\text {max }}^{(n-1)}$ ) (not between $\tilde{P}_{n-1, m i n}, \widetilde{P}_{n-1, \text { max }}$ ) and the verification of $/ 2.18 /$ is not neoessary. only after it $\varphi_{1}, \ldots, \varphi_{n-1}$ are drawn in $(0,2 \mathcal{I}), \cos \theta_{n-1}$ is calculated by $/ 2.14 /$ and $\Phi\left(P_{1}, \ldots, \varphi_{n-i}\right)$ by /2.I6/.

The realization of the density distribution using the rejection technique has in this case some pecularities. It $1 s$ impossible to appiy this technique directly since $\phi_{\text {mac }}=\infty$ for $P_{n-1}=0$ if $F$ is innite everywhere. But just at $P_{n-1} \rightarrow 0$ the interval of the change $P_{n-1}(2.19)$ contracts into a point, 1.0. $P_{n-1}$ need not be drawn. Let the accuracy for setting the momenta in Table be $\varepsilon$, then it is possible to find $P_{\text {min }}^{(n-1)}$ so small that the interval of the change $P_{n-1}$ will be less than $\varepsilon$. , and for $P_{n-1} \leqslant P_{\min }^{(n-1)}$ 1t may be supposed that $P_{n-1}=P_{n-1}^{*}$. If $P_{n-1}>p_{m i n}^{(n-1)}$ then the distribution is realized by a rejection technique, and the maximum / $2.16 /$ is taken as a maximum of $\phi$ with the change of $P_{n-1} f o r P_{m i n}^{(n-1)}$. If a preliminary evaluation of $\phi_{m a x}$ is too overestimated or impossible due to a oomplex form of $F$, it may be improved in the course of calculation. Having calculated $\phi=\phi^{\prime \prime}, \ldots, \phi^{(\lambda)}$ for a sufficiently large number $\nu$ of sets of the components $\overrightarrow{P_{1}}, \ldots, \vec{P}_{n-1}$ one may assume the greatest value of $\boldsymbol{p}$ on the obtained assembly of sets as $\phi_{\text {max }}$ : After that having ohoosen $\nu$ random numbers $\alpha^{(1)}, \ldots, \alpha^{(\nu)}$ on ( $0, \phi_{\text {max }}$ ) and comparing $\phi^{(1)}, \ldots, \phi^{(y)}$ with them, it is necessary to reject those sets for which $\alpha>\Phi$.

After the distribution of the density $D$ has been realized the tabulation is completed by the calculation of $\vec{P}_{n}=-\vec{P}_{n-1}-\vec{P}_{n-1}$ and by transforming of all the momenta into the system OXYZ by /2.9/-/2.12/.

The Fermi Model $F \equiv I$ enables to simplify a little the calculation method. In this case one may in / 2.I3/ integrate over $\varphi_{n-1}$ and over $\rho_{n-1}$ (more exactly, over $e_{n-1}$ within the limits

$$
\left.e_{\max }=\frac{E_{n-1}^{*} E_{n-1} \mp P_{n-1}^{*} P_{n-1}}{M_{n-1}}\right)
$$

and to obtain the density distribution over $\vec{P}_{1}, \ldots, \vec{P}_{n-2}$ :

$$
\phi^{1}\left(\overrightarrow{P_{1}}, \ldots, \overrightarrow{P_{n-2}}\right)=\frac{4 \pi}{3} P_{n-1}^{*} \frac{3 E_{n-1}^{*} E_{n}^{*} E_{n-1}^{2}-P_{n-1}^{* 2} P_{n-1}^{2}}{M_{n-1}^{3}} \cdot \prod_{i=1}^{n-2} p_{i}^{2}
$$

Here the sequence of samplings is the same as in the general case, but comes to an end on the

I/ If $n=3, / 2.25 /$ coincides $w i t h / 4 /$ from $/ 6 /$.
sampling $P_{n-2}, \cos \theta_{n-2}$. On obtaining the points uniformly distributed in $D_{n-I}$ the nonuniform distribution of the density $\phi^{\prime}$ is being realized using the "rejection technique". $\phi_{\text {max }}^{\prime}$ may be calculated beforehand, and, namely, it may be shown that the greatest value of中' is realized when $\overrightarrow{P_{n-I}}=0$, and the momenta $\overrightarrow{P_{1}}, \ldots, \vec{P}_{n-2}$ satisfy the equation system

$$
\frac{P_{1}^{2}}{e_{1}}=\frac{P_{2}^{2}}{e_{2}}=\ldots=\frac{P_{n-2}^{2}}{e_{n-2}}=\frac{2 P_{n-1}^{* 2} E_{n-1}^{*} E_{n}^{*} E_{n-1}}{\left(E_{n-1}^{*} E_{n}^{*}\right)^{2+} P_{n-1}^{* 2}\left(E_{n-1}^{* 2}-E_{n-1}^{*} E_{n}^{*}+E_{n}^{* 2}\right)}
$$

This system is easily solved using the iteration method.
Further it is necessary to select the momentum $P_{n-1}$. Instead of this it is better to choose the energy $e_{n-1}$ using the direct method $/ B /$, i.e., to solve the aquation

$$
2 e_{n-1}^{3}-3 E_{n-1} e_{n-1}^{2}=\left(2 e_{\min }^{3}-3 E_{n-1} e_{\min }^{2}\right)(1-\alpha)+\left(2 e_{\max }^{3}-3 E_{n-1} e_{\max }^{2}\right) \alpha_{, / 2} .27 /
$$

Where $\alpha \quad$ is uniformly distributed in $/ 0, I /$.
only after this $1 t$ is meaningful to pick $\varphi_{1}, \ldots, \varphi_{n-1}$.
The characteristic feature of the method stated in § I-2 is the presence of the unsuccessful samplings, i.e., the realization of the nonuniform distributions using the rejeotion teohnique. Evidently, when the number of unsuccessful samplings is great the method becomes practically useless for tabulating. One fails to estimate the efficiency of the method In the general form. However, the numerical experiments has shown that for the reaotions on meson production by protons at $E=4.72 \mathrm{BeV}$ the efficiency of the method is $50 \%$./the production of one pion / and $10 \% /$ two pions/. When the number of mesons is great the efficiency is too small to apply the method described above in the hand calculations. Still it is possible to make up the tables when $3-4$ mesons are produced by the hand caloulation if $\phi^{\prime}$ is integrated some times more. The use of electronic computer will enable to make up large tables of 5-6 -particle reactions applying the method described here.

## 3. TABLE OF RANDOM STARS

The table of $n$-particle reactions is difficult to compare with experimental data since neutral particles fail to be detected in the experiment. It must be incorporated into the table of random stars created in tho collision of the two given particles at the given energy E.

To make up such a table it is necessary to consider all kinds of reactions having an . appreciable weight. The magnitudes of the statistioal weights define the average number of lines for each reaction.

In the Fermi model the statistical weights may be oalculated and are already known for many reaotions. But for the simulating of the interaction with an arbitrary square of the
matrix element the problem of the calculation of the statistical weight may appear to be as difficult as the tabulation of random stars. However, the above method for the making up of the table of n-particle reactions enables in principle to determine the phase volume $W$ the main part of the expression for the statistical weight. Indeed, it is clear that the phase volume $W$ is equal to the product of the volume of the region $\mathscr{D}$ n an expected value $\infty / 2 . I 6 /$ over this region. The volume of the region $\mathscr{D} n$ is equal to the fraction of the successful hits of the sets $\overrightarrow{P_{1}}, \ldots, \overrightarrow{P_{n-1}} \quad$ to this region multiplied by the volume of the rectangle with the known dimensions circumscribed around $G D n$. The appiioation of this method for calculating the phase volumes would require a considerable number of samplings due to great spread in the values of $\Phi$, , the smaller the phase volume the greater the spread.

If the statistical weights of the most important reactions occuring at the given energy are known then the making up of the table of random stars is not difficult. Let us arrange all the reactions $\mathcal{N}_{1}, \mathcal{N}_{2}, \ldots, \mathcal{N}_{j}, \ldots$ in an arbitrary order and give its own interval $\delta_{\mathcal{N}_{j}}$ to each of them on the segment ( $O, I$ ). The length of the interval must be proportional to the statistical weight of the reaction $\mathcal{N}_{j}$. Then for drawing the next line of the table one throws randomly a point on the segment $(O, I)$. The number of the interval $\delta_{\mathcal{N} j}$ to which it will be incident points out the reaction which will be represented in this line. Besides it is being cleared up by sampling which of the momenta in this line should be considered belonging to the oharged particles. The momenta of other particles as unobservable need not be introduced into the line.

In the table made up in such a way $a /$ the reactions will be shuffled randomly b/ the number of lines reffering to each reaction will fluctuate with respect to that expected by the statistical weight. The law of these fluctuations will be the same as in the observation of the real stars. It should be noted in order to increase the accuracy of small tables the fluctuations may be forbidden.

Using the computer giving the stars with $\quad n=6 \quad$ by the above method a majority of reactions induced by protons of the kinetic energy $I O \mathrm{BeV}$ may be taken into account. The stars with different number of rays will be present.

The accuracy of the results obtained from the table of random stars is mainly determined by the same factors as in case of the real stars. There are no errors in the determination of energy, angles of particie emergence; one may neglect oscillations of the frequency of some reactions with respect to their weight. However, the results will be influenced by the error in the calculation of the statistical weights of all reactions. When the errors in the weights are of the order of $10 \%$ the making up of tables of mare than 1000 lines is
hardly worth while. The surplus of accuracy which is taken is necessary to obtain the conditional distributions and for the comparison of the distributions at different energies. The error in the statistical weights at different $E$ is levelled since they are more or less systematical. Therefore, the results of the comparison of the distributions at different $E$ may have a greater accuracy.

An interesting possibility of increasing the accuracy appears in the tabulation of random stars in the laboratory system of reference. Then there is no neoessity any more to transform real stars into the centre of mass system - this problem is solved at high energies rather roughly. Thus, the analysis of the experimental data becomes easier and the accuracy of comparing the table of random stars with the experimental data increases. The formulae for the calculations in both systems are identical as concerns their complexity /in §2 everywhere $\vec{P} \neq 0 \%$ Practically, however, at $\vec{P} \neq 0$ the interval of the allowed values of the momenta reduces/narrow beamsl/ and the interval of the allowed values of the momenta $\vec{p}_{K}$ extends. Both these factors reduce the efficiency of the samplings $\overrightarrow{p_{k}} \quad$. Therefore, it is better to transform the momenta in the table of stars in the centre of mass system into the laboratory system.

## CONCLUSION

It is the purpose of this paper to clear up the possibility of making up the table of random stars. A way for it is given here. Evidently, in the first turn one will succeed in reproducing the Fermi model and that of isobar in the form of a table. One may hope that the reproduction of the reactions with 6-7 secondary particles for the modern computers is quite real. It means that it is possible to make up for instance, a table of stars created p-p collisions with the $I O$ BeV accelerator.

The way of making up such a table for these models is such that the distribution of the reaction products over the angles is isotropic. To explain the nonisotropy really observed one has to choose a definite suitable dependence of $\mathcal{F}$ upon momenta and angles.

The described method of making up the table of random stars is accepted in different kinds of the matrix element square $F$. The complexity of the function $\mathcal{F}$ is of no importance since a major period of time during the work of the computer has to be spent far rejecting the unsuccessiful momentum componenta, but not for caloulating the aocepted momenta. Therefore there is a possibilits of finding a satisfactory $\mathcal{F}$ from a phenomenologically standpoint form of interaction by testing different ideas about the interaction and comparing the obtained statistical weights, distributions, correlations with the experimental data Such a method might have been applied forstudying the decay of unstable particles.

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Fig.I


Fig. 2

- I7 -


Fig. 4
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