

# сообиепй oбheanisnine <br> nietutyta hatpliax исенеднанй аупиа 

E2-87-78

## B.Z.lliev

THE DEVIATION EQUATION AS AN EQUATION OF MOTION

## 1. INTRODUCTION

In this work we are going to look on the generalized deviation equation from a dynamical point of view. This is very natural because of its definition (see $13 \%$, sect. 1.3 and ${ }^{\prime 4 /}$, sect. IV.3) because it is intuitively clear that the equations (laws) of the (relative) motion of two point particles in spaces $L_{n}$ with the affine connection could be considered also as deviation equations.

Ideas analogous to ours have been considered in more special cases in ${ }^{\prime, 2,5,6 /}$. In $^{\prime 6}$ the deviation equation is derived : f two infinitesimally near interacting point particles in the space-time of general relativity. In 's' the influence of small perturbative forces of dissipative and periodical kinds in the right-hand side of the usual equation of geodesic deviation in a Riemannian space $V_{n}(n=4)$ is treated. Paper is an investigation of the relative dynamics of geodesies in spaces with an affine connection without torsion in terms of absolute two-point derivatives of a geodesic radius-vector. The closest to our work is ${ }^{\prime \prime}$ ', where exact and approximate equations of relative motion of trial punt particles in the field of external forces in a Riemannia space are obtained in terms of absolute two-point derivatives of a geodesic radiusvector.

The purpose of the present work, in which the notation is the same as in $13.4 \%$, is to show that in the general case the equation of motion of two point particles in the field of external force in space with the affine connection is a special case of the generalized deviation equation. After some preliminary considerations and one simple example (sections 2 and 3) we derive in section 4 in the general case the deviation equation in the form of an equation of motion. In sections 5 and 6 we consider two examples illustrating the developed general theory.

## 2. THE BASIC PRELIMINARY CONSTRUCTION

Let us define the curve $x:\left[s^{\prime}, s^{\prime \prime}\right] \rightarrow L_{n}, L_{n}$ is the space with the affine connection as unique solution of the following ini-tial-value problem:

$$
\begin{equation*}
\frac{\mathrm{Du}(\mathrm{~s})}{\mathrm{ds}}=F(\mathrm{~s}, \mathrm{x}(\mathrm{~s}), \mathrm{u}(\mathrm{~s})), \quad \mathrm{u}^{\alpha}:=\mathrm{u}^{\alpha}(\mathrm{s}):=\frac{\mathrm{d} \mathrm{x}^{\alpha}(\mathrm{s})}{\mathrm{ds}} \tag{2.1}
\end{equation*}
$$

$x\left(s_{0}\right)=x_{0} \in L_{n}, u\left(s_{0}\right)=u_{0} \in T_{x_{0}}\left(L_{n}\right), s, s_{0} \in\left[s^{\prime}, s^{\prime \prime}\right]$,
where $D / d s$ is the covariant derivative with respect to $s$ along $x(s)$, and $F$ is a continuous function of its arguments.

Physically, we shall interpret $x$ as a trajectory (worlri iine) of an observer who is affected by the force field $F$ having a meaning of a force per unit mass and whose trajectory passes through the point $x_{0}$ with a velocity $u_{0}$.

Let us define the family of curves $\gamma_{\mathrm{s}}:\left[\mathrm{r}^{\prime}, \mathrm{r}{ }^{\prime \prime}\right] \rightarrow \mathrm{L}_{\mathrm{n}}, \mathrm{s} \in$ $\epsilon\left[s^{\prime}, s^{\prime \prime}\right]$ as the unique solution of the following initial-value problem:
$\left.\frac{\mathrm{D}}{\mathrm{ds}}\right|_{\gamma_{\mathrm{s}}}\left(\gamma_{\mathrm{s}}^{\prime}(\mathrm{r})\right)=\mathrm{F}_{\mathrm{s}}(\mathrm{r}):=\mathrm{F}_{\mathrm{s}}\left(\mathrm{r}, \gamma_{\mathrm{s}}(\mathrm{r}), \gamma_{\mathrm{s}}^{\prime}(\mathrm{r})\right) \subseteq \mathrm{T}_{\gamma_{\mathrm{s}}(\mathrm{r})}\left(\mathrm{L}_{\mathrm{B}}\right)$, $\gamma_{\mathrm{s}}^{\prime a}(\mathrm{r}):=\partial \gamma_{\mathrm{s}}^{a}(\mathrm{r}) / \partial \mathrm{s}$,
$\left.\gamma_{\mathrm{s}}(\mathrm{r})\right|_{\mathrm{s}=\mathrm{s}_{\mathrm{o}}}=\chi(\mathrm{r}) \in \mathrm{L}_{\mathrm{n}},\left.\gamma_{\mathrm{s}}^{\prime}(\mathrm{r})\right|_{\mathrm{s}=\mathrm{s}_{\mathrm{o}}}=\chi^{\prime}(\mathrm{r}) \in \mathrm{T}_{\chi(\mathrm{r})}\left(\mathrm{L}_{\mathrm{n}}\right)$,
where $F_{s}, \chi$ and $\chi^{\prime}$ are given smooth functions of their arguments, $r \in\left[r^{\prime}, r^{\prime \prime}\right], s \in\left[s^{\prime}, s^{\prime \prime}\right]$, the number $s_{o} \in\left[s^{\prime}, s^{\prime \prime}\right]$ is the same as in (2.2) and $\mathrm{D} / \mathrm{ds} \mid \gamma_{\mathrm{s}}$ means the covariant derivative with respect to $s$ along the $r$-curve $\gamma_{s}(r), r=$ const (by $D / d s$ we denote the covariant derivative with respect to $s$ along the curve $x(s)$, cf. (2.1)).

Let the curves $x_{a}:\left[s_{a}^{\prime}, s_{a}^{\prime \prime}\right] \rightarrow L_{n}, a=1,2$ be defined by $x_{1}\left(\tau_{1}(s)\right):=\gamma_{s}\left(r^{\prime}\right), x_{2}\left(\tau_{2}(s)\right):=\gamma_{s}\left(r^{\prime \prime}\right)$,
where $s \in\left[s^{\prime}, s^{\prime \prime}\right]$ and $r_{a}:\left[s^{\prime}, s^{\prime \prime}\right] \longrightarrow\left[s_{a}^{\prime}, s_{a}^{\prime \prime}\right], a=1,2$ are some given smooth mapping. So, due to (2.4) the curves $x_{1}$ and $\mathrm{x}_{2}$ pass through the points $\chi\left(\mathrm{r}^{\prime}\right)$ and $\chi\left(\mathrm{r}^{\prime \prime}\right)$, respectively, and have at them tangent vectors $\chi^{\prime \prime}\left(r^{\prime}\right)$ and $\chi^{\prime \prime}\left(r^{\prime \prime}\right)$, respectively.

Physically, we shall interpret $\mathrm{F}_{\mathrm{s}}$ as a force (per unit mass) acting in the 2-dimensional submanifold (surface) $\left|\gamma_{\mathrm{s}}(\mathrm{r})\right| \mathrm{s} \in$ $\left.\in\left[s^{\prime}, s^{\prime \prime}\right], r \in\left[r^{\prime}, r^{\prime \prime}\right]\right\}$. (It is clear that $F_{s}$ and $F$ are analogues of the Minkowski four-force per unit mass). The curves $x_{1}$ and $x_{2}$ shall be considered as trajectories (world lines) of point particles observed from the observer defined above whose trajectory is defined by (2.1) and (2.2).

At the end, as has been done in '4', we shall define a family of curves $\eta_{\mathrm{S}}:\left[\rho^{\prime}, \rho^{\prime \prime}\right] \rightarrow \mathrm{L}_{\mathrm{n}}$, such that $\eta_{\mathrm{s}}\left(\rho^{\prime}\right):=\mathrm{x}_{1}\left(\tau_{1}(\mathrm{~s})\right)$ and $\left.\eta_{\mathrm{s}}\left(\rho^{\prime \prime}\right):=\mathrm{x}(\mathrm{s})\right), \mathrm{s} \in\left[\mathrm{s}^{\prime} \mathrm{s}^{\prime \prime}\right]$.

In the next sections we shall consider the problem of finding the deviation equation of $x_{2}$ with respect to $x_{1}$ as it is observed from $x$ at some point $x(s)$ (for the corresponding definitions see ${ }^{3.4 \prime}$ ). This means to express the relative (deviation) acceleration $D^{2} h_{12}(s, x) / d s^{2} \quad\left(h_{12}(s, x)\right.$ is the corresponding deviation vector) through the defined above quantities, and first of all, through the force field $\mathrm{F}_{\mathrm{s}}$, which practically means to write down the equation of relative motion of $x_{2}$ with respect to $x_{1}$ relatively to $x$. In this sense, the equation of motion is a special case of the generalized deviation equation.
3. THE deviation equation as an equation of motion: EUCLIDEAN CASE

To derive the deviation equation we have to do the following: using the definition of the deviation vector (see ${ }^{\prime 4}$, sect. IV. I), we have to compute its second covariant derivative $\mathrm{D}^{2} \mathrm{~h}_{12}(\mathrm{~s}, \mathrm{x}) / \mathrm{ds}^{2}$ (along $\mathrm{x}(\mathrm{s})$ ) and next to substitute in the obtained expression the formulae (2.1)-(2.4).

To make these ideas more clear we shall consider at first the most simple possible case, which we shall call an "Euclidean case": let us take $L_{n}$ to be an $n$-dimensional Euclidean space $E_{n}$ and $I^{\gamma}$ to be a parallel transport along $y$. Then (see $/{ }_{4}{ }^{n}$, sect. III.3.1 and $14 /$, sect. IV.1) we find the deviation vector $h_{12}(s, x)$ to he with the components

$$
\begin{equation*}
h_{12}^{a}(s, x)=x_{2}^{a}\left(r_{2}(s)\right)-x_{1}^{\alpha}\left(r_{1}(s)\right), \tag{3.1}
\end{equation*}
$$

so the deviation equation is simply

$$
\begin{equation*}
\frac{D^{2} h_{12}(s, x)}{d s^{2}}=F_{2}-F_{1} \text {. } \tag{3.2}
\end{equation*}
$$

where
$\mathrm{F}_{1}:=\mathrm{F}_{\mathrm{s}}\left(\mathrm{r}^{\circ}, \mathrm{x}_{1}\left(r_{1}(\mathrm{~s})\right), d \mathrm{x}_{1}\left(\mathrm{r}_{1}(\mathrm{~s})\right) / \mathrm{ds}\right)$,
$\mathrm{F}_{2}:=\mathrm{F}_{\mathrm{s}}\left(\mathrm{r}^{\prime \prime}, \mathrm{x}_{2}\left(\mathrm{r}_{2}(\mathrm{~s})\right), \mathrm{d} \mathrm{x}_{2}\left(\mathrm{r}_{2}(\mathrm{~s})\right) / \mathrm{ds}\right)$.
are the forces per unit mass acting on the particles 1 and 2 , respectively.

Evidently, eq. (3.2) is nothing else but the second Newton law of dynamics, i.e. the equation of the relative motion of the particle 2 with respect to the particle 1 in the force field $F_{s}$.
4. THE DEVIATION EQUATION AS AN EQUATION OF MOTION: GENERAL CASE

We shall begin this section with a mathematical definition necessary for the following considerations:

Let $A^{i_{1} \ldots i_{p}}{ }_{j} \ldots j_{q}\left(z_{1}(s), \ldots, z_{p+q}(s)\right)$ be components of the ( $p+q$ )-point $C^{1}$-tensor $A$ from the space
$T_{z_{1}(s)}\left(L_{n}\right) \otimes \cdots T_{z_{p}}(s)\left(L_{n}\right) \otimes T_{z_{p+1}}^{*}(s)\left(L_{n}\right) \otimes \ldots \omega_{z_{p}+q^{(s)}}\left(L_{n}\right)$, where $z_{a}:\left[s^{\prime}, s^{\prime \prime}\right] \rightarrow L_{n}, a=1, \ldots, p+q$ are some $C^{1}$-maps and $s \in\left[s^{\circ}, s^{\circ}\right]$. Then, by definition
$\frac{D}{d s} A^{i_{1} \cdots i_{p}} j_{1} \ldots j_{q}\left(z_{1}(s), \ldots, z_{p+q}(s)\right):=$
$=\frac{d}{d s} A^{i_{1} \ldots i_{p}} j_{1} \ldots j_{q}\left(z_{1}(s), \ldots, z_{p+q}(s)\right)+\sum_{b=1}^{p} \Gamma^{i_{b}} \cdot k p\left(z_{b}(s)\right) \times$
$\left.\times A^{i_{1} \ldots i_{b-1} k i_{b+1} \cdots i_{p}} j_{1 \cdots j_{q}}\left(z_{1}(s), \ldots, z_{p+q^{\prime}}(s)\right) z_{b}^{\prime l}(s)-\sum_{b=1}^{q} \Gamma^{k} \cdot j_{b} p^{\left(z_{p+b}\right.}(s)\right) \times$
$\times A^{i_{1} \cdots i_{p}} j_{1} \ldots j_{b-1} k j_{b+1} \ldots j_{q}\left(z_{1}(s), \ldots, z_{p+q}(s)\right) z_{p+b}^{\ell}(s)$,
where $z_{a}^{a}(s):=d z_{a}^{\alpha}(s) / d s \quad, a=1, \ldots, p+q$ and $\Gamma^{i} \cdot j k$ (y) are the coefficients of tie affine connection at $y \in L_{n}$.

It is not difficult to see that (4.1) are components of a many-point tensor of the same tensor space as tre initial tensor A.

From this definition it is easy to check that the operator $\mathrm{D} / \mathrm{d}$ s is a differentiation with respect to the tensor products
of (many-points) tensors, i.e. the usual Liebniz rule holds for tensor products $(D(A * B) / d s=(D A / d s) \otimes B+A \&(D B / d s))$ and $D / d s$ commutes with the contraction operator, and as a consequnce of this, $D / d s$ commutes also wh the contracted tensor products (e.g.,

$$
\begin{aligned}
D\left(A _ { \cdot j } ^ { j } \left(z_{1}(s),\right.\right. & \left.z_{2}(s)\right) B^{j}\left(z_{2}(s)\right) / d s= \\
& =\sum_{j=1}^{n}\left\{\left.\left[D\left(A^{j} \cdot j\left(z_{1}(s), z_{2}(s)\right) B^{k}\left(z_{3}(s)\right)\right) / d s\right]\right|_{\substack{z_{3}=z_{2} \\
k=j}}\right) .
\end{aligned}
$$

Let us now recall the definition of the deviation vector $h_{12}(s, x)$ of $x_{2}$ with respect to $x_{1}$ relatively to $x$ at the point $x(s)\left(\right.$ see ${ }^{4,}$, sect. IV.1):
$h_{12}(s, x):=I_{x_{1}\left(t_{1}(s)\right)+x(s)}^{\eta_{\mathrm{s}}} \int_{r^{\prime \prime}}^{r^{\prime \prime}} \mathrm{I}_{\gamma_{s}(r) \rightarrow x_{1}\left(r_{1}(s)\right)}^{\gamma_{s}} \quad \gamma_{s}(r) d r$,
or $\left(\operatorname{see}^{1 / 3 *}\right.$, sect. II.2.4)
$h_{12}^{k}(s, x)=H^{k} \cdot p \quad \int_{r^{\prime}}^{\prime \prime} A^{p} \cdot \dot{\gamma}_{s}^{q}(r) d r$,
where $\gamma_{s}^{*}(r):=\partial y_{s}^{a}(r) / \partial r \quad, I \eta_{s} \quad$ and $I^{\gamma_{s}} \ldots$ are the generalized transports along $\eta_{s}$ and $\gamma_{s}$, respectively, defined in $3^{\prime}$, sect. II. 1 and
$H^{k} \cdot p:=H^{k} \cdot p\left(x(s), \gamma_{s}\left(r^{\prime}\right)\right):=\left(I_{\gamma_{s}}^{\eta_{s}}\left(r^{\prime}\right) \rightarrow x(s)^{k} \cdot p\right.$,
$\left(\mathrm{H}^{-1}\right)_{\cdot q}^{p}:=H_{\cdot q}^{p}\left(\gamma_{\mathrm{s}}\left(\mathrm{r}^{*}\right), \mathrm{x}(\mathrm{s})\right)=\left[\left(\mathrm{I}_{\gamma_{\mathrm{s}}}^{\eta_{\mathrm{s}}}\left(\mathrm{r}^{*}\right) \cdot \mathrm{x}(\mathrm{s})^{-1}\right]_{\cdot q}^{\mathrm{p}}\right.$,
$\Lambda_{\cdot q}^{p}:=\Lambda_{\cdot q}^{p}\left(\gamma_{s}\left(r^{*}\right), \gamma_{s}(r)\right):=\left(I_{\gamma_{s}}^{\gamma_{s}}(r) \rightarrow \gamma_{s}\left(r^{\prime}\right)^{p} \cdot q^{p}\right.$.
are the components of the two-point tensors representing the corresponding linear mapping $I .$. in a given basis (see $3^{\prime \prime}$, sect. II.2.4).

Now we are ready to derive the deviation equation of $x_{2}$ with respect to $x_{1}$ relatively to $x$ at $x(s)$ :

For some time, for the sake of shortness, we shall suppress the indices and the arguments of all quantities. Thus, we can rewrite (4.3) symbolically in the form
$h=H \cdot \int \Lambda \cdot \dot{\gamma}$,
where the dot (•) means a contracted tensor product (e.g., $\Lambda \cdot \dot{\gamma}$ is a vector with components $(\lambda \cdot \dot{y})^{j}:=\Lambda j_{k} \dot{\gamma}^{k}=\Lambda_{\cdot k}^{j}\left(\gamma_{s}\left(r^{\prime}\right)\right.$, $\left.y_{\mathrm{s}}(\mathrm{r})\right) \dot{y}_{\mathrm{s}}^{\mathrm{k}}(\mathrm{r}), \int \mathrm{A} \cdot \dot{\gamma}:=\hat{i} \cdot \dot{\gamma} \mathrm{dr}$ and so on).

So, using (4.1) and (4.5), we get:
$\frac{D^{2}}{d s^{2}}!x:=\frac{D}{d s} \cdot\left(\frac{D}{d s} ; h\right)=\frac{D^{2} H}{d s^{2}} \cdot H^{-1} \cdot h+$
$+2 \frac{\mathrm{DH}}{d \mathrm{~s}} \cdot \int\left(\frac{\mathrm{D} \Lambda}{\mathrm{ds}} \cdot \dot{\gamma}-\Lambda \cdot \frac{\mathrm{D} \dot{\gamma}}{\mathrm{d} s}\right)=H \cdot \int\left(\frac{\mathrm{D}^{2} \Lambda}{d s^{2}} \cdot \dot{\gamma}+2 \frac{\mathrm{D} \Lambda}{\mathrm{ds}} \cdot \frac{\mathrm{D} \dot{\gamma}}{\mathrm{ds}}+\Lambda \cdot \frac{\mathrm{D}^{2} \dot{\gamma}}{\mathrm{~d} \mathrm{~s}^{2}}\right)$.
(4.6)

By a direct calculation one can check the following equality:
$\frac{D^{2} j^{p}}{d s^{2}}=\left\langle R_{\cdot j j k}^{p} \gamma^{1} \cdot i-T_{\cdot j k}^{p} \frac{D y^{\prime j}}{d s}+\frac{D T^{p} \cdot j k}{d s} y^{\cdot j \gamma^{k}}+\right.$

where $R_{\cdot i j k}^{p}$ and $T_{. j k}^{P}$ are the components of the curvacure and
 I. $\alpha y, \delta:=2 \Gamma^{\alpha} \beta y$ x .

It is not difficult to show that the sum in square orackets in $(4.7)$ is

$$
\begin{equation*}
\left(\cdots \left\lvert\,=\frac{D}{d s}\left(\frac{D y^{-}}{d s}\right)+T{ }^{p} \cdot y^{1} \cdot \frac{D i^{k}}{d s}\right.\right. \tag{4.8}
\end{equation*}
$$

Substituting (4.8) into (4.7), then the so-obtained equality in (4.6) and using (2.1)-(2.4), we finally get the following deviation equation:
$\frac{D^{2}}{d s^{2}} x_{12}^{k}(s, x)=\frac{D^{2} H^{k} \cdot p}{d s^{2}}\left(H^{-1}\right)_{\cdot q}^{p} h_{1 g}^{q}(s, x)+$
$+2 \frac{D H_{p}^{k}}{d s} \int_{r^{*}}^{p^{\prime \prime}} d r\left(\frac{D X_{q}^{p}}{d s} \dot{y}_{s}^{q}(r)+A_{q}^{p} \frac{D \dot{\gamma}_{s}^{q}(r)}{d s}\right)+$
$+H_{* p}^{k} \int_{r^{\prime}}^{r^{\prime \prime}} d r\left\{\frac{D^{2} A_{r q}^{p}}{d s^{2}} ; s^{q}(r)+2 \frac{D A_{\cdot q}^{p}}{d s} \frac{D \gamma_{s}^{q}(r)}{d s}+\right.$
$+\Lambda_{\cdot q}^{p}\left[\left(R_{\cdot i j \ell}^{q}\left(\gamma_{S}(r)\right) \gamma_{s}^{\prime j}(r) \gamma_{s}^{\prime j}(r)+T_{\cdot j \ell}^{q} \ell\left(\gamma_{S}(r)\right) F_{s}^{j}(r)+\right.\right.$
$\left.+\frac{D T{ }^{q}{ }^{j \ell}\left(\gamma_{s}(r)\right)}{d s} \cdot y_{s}^{\prime j}(r)\right) \dot{y}_{s}^{\prime \ell}(r)+$
$\left.\left.+T_{\cdot j \ell}^{q}\left(\gamma_{s}(r)\right) \gamma_{s}^{\prime j}(r) \frac{D \dot{\gamma}_{s}^{\prime}(r)}{d s}+\frac{D F_{s}^{q}(r)}{d r} \right\rvert\,\right\}$.

We want to remark that the force $F$ acting on the observer is also presented in (4.9), but implicitly it is hidden in the second derivative $\mathrm{D}^{2} \mathrm{H}^{k} \cdot \mathrm{p} / \mathrm{d} \mathrm{s}^{2}$.

Thus, we see that the deviation equation (4.9) really has a meaning of an equation of motion of the particle 2 with respect to the particle 1 as it is observed from an observer with a trajectory $x$.

From a dynamical point of view, the most important terms in the deviation equation (4.9) are those in which the force $F_{S}(r)$ appears, viz.
$H^{k}{\underset{p}{p}}_{r^{\prime \prime}}^{r^{\prime}} \Lambda_{\cdot q}^{p}\left(T_{\cdot j p}^{q}\left(y_{s}(r)\right) F_{s}^{j}(r) \gamma_{s}^{p}(r)+\operatorname{DF}_{s}^{q}(r) ; d r\right) d r=$
$\cdot\left[H^{k} \cdot p\left(\Lambda_{\cdot j}^{p}\left(y_{s}\left(r^{\prime}\right), \gamma_{s}\left(r^{\prime \prime}\right)\right) F_{s}^{j}\left(r^{\prime \prime}\right)-F_{s}^{p}\left(r^{\prime}\right)\right)\right]+$
$+H_{\cdot p}^{k} \stackrel{r^{\prime \prime}}{r^{\prime}}\left(A_{\cdot q}^{p} T_{\cdot j \ell}^{q} \ell\left(\gamma_{s}(r)\right) \gamma_{s}^{\ell}(r)-\frac{D \Lambda_{\cdot j}^{p}\left(\gamma_{s}\left(r^{\prime}\right), \gamma_{s}(r)\right)}{d r}\right) \mathrm{F}_{s}^{j}(r) d r$,
where we have done an evident integration by parts of the term $\Lambda_{\text {p }}^{p} \mathrm{DFs}_{\mathrm{g}}^{\mathrm{g}}(\mathrm{r}) / \mathrm{dr}$

From (4.10) we see that the only independent of $y$ terms in (4.9) are those in the square brackets in (4.10), which are equal to the definec by means of the generalized transports IYs. and $1 \eta$. difference of the forces acting on the observed particles.

## 5. EXAMPLE: EUCLIDEAN CASE

The most simple application of the general theory from the previous section is the derivation of the Euclidean equation of motion (3.2).

In the Euclidean case (see section 3 ), we replace $L_{n}$ with the $\mathrm{E}_{\mathrm{n}}$ space and $\mathrm{I}^{\gamma}$ with a parallel transport along $y$. So, working in a global coordinate basis in which $\Gamma . \beta \gamma \equiv 0$, we have $H^{a} \cdot \beta=\Lambda^{a} \cdot \beta=\delta_{\beta}^{a}\left(\delta_{\beta}^{a}=1\right.$ for $a=\beta$ and $\delta_{\beta}^{a}=0$ for $a \neq \beta$ ). Thus, in this case due to (4.9) the deviation equation is
$\frac{D^{2}}{d s^{2}} f_{x} h_{12}(\mathrm{~s}, \mathrm{x})=0+0+\int_{\mathrm{r}^{\prime}}^{\mathrm{r}^{\prime \prime}} \mathrm{dr}\left(0+\cdots 0+\frac{\mathrm{d} \mathrm{F}_{\mathrm{s}}(\mathrm{r})}{\mathrm{dr}}\right)=\mathrm{F}_{\mathrm{s}}\left(\mathrm{r}^{\prime \prime}\right)-\mathrm{F}_{\mathrm{s}}\left(\mathrm{r}^{\prime}\right)$, which, because of (2.3) and (3.3), is identical with (3.2).
6. EXAMPLE: THE FIRST ORDER DEVIATION EQUATION (EQUATION OF MOTION)

In this section we shall derive the first approximation to the exact deviation equation (equation of motion) (4.9), or more precisely, we shall write down eq. (4.9) within terms of an order of $O\left(\rho^{\prime \prime}-\rho^{\prime}\right)$ and $O\left(\left(r^{\prime \prime}-r^{\prime}\right)^{2}\right)$ with respect to the parameters of the families $\eta_{\mathrm{s}}:\left[\rho^{\prime}, \rho^{\prime \prime}\right] \longrightarrow \mathrm{L}_{\mathrm{n}}$ and $\gamma_{\mathrm{s}}$ : $:\left[r^{\prime}, r^{\prime \prime}\right] \rightarrow L_{n}$, respectively (see section 2). As we shall see below, this approximation is independent of the concrete choice of the generalized transports $I \eta_{\mathrm{s}}$ and $\mathrm{I}^{\gamma \mathrm{s}}$.

Having in mind that by definition (see section 2) $\eta_{\mathrm{s}}\left(\rho^{\prime}\right):=$ $:=\mathrm{x}_{1}\left(\tau_{1}(\mathrm{~s})\right):=\gamma_{\mathrm{S}}\left(\mathrm{r}^{\circ}\right), \quad \eta_{\mathrm{s}}\left(\rho^{\prime \prime}\right):=\mathrm{x}(\mathrm{s}), \gamma_{\mathrm{s}}\left(\mathrm{r}^{\prime \prime}\right)=: \mathrm{x}_{2}\left(\tau_{2}(\mathrm{~s})\right)$ and $r \in\left[r^{\prime}, r^{\prime \prime}\right]$ and using eq. (2.6) from ${ }^{\prime \prime} \rho^{\prime}$, sect. II. 2.4 (which means that $\mathrm{I}_{\mathrm{y} \rightarrow \mathrm{y}}:=$ id for any $\mathrm{y} \in \mathrm{L}_{\mathrm{n}} \rightarrow \mathrm{see}^{\prime / 3 /}$, sect.II.1), we immediately derive from (4.4)
$\mathrm{H}_{\cdot \mathrm{p}}^{\mathrm{k}}=\delta_{\mathrm{p}}^{\mathrm{k}}+\mathrm{O}\left(\rho^{\prime \prime}-\rho^{\prime}\right), \quad\left(\mathrm{H}^{-1}\right)_{\cdot \mathrm{q}}^{\mathrm{p}}=\delta_{\mathrm{q}}^{\mathrm{p}}+\mathrm{O}\left(\rho^{\prime \prime}-\rho^{\prime}\right)$,
$\Lambda_{\cdot q}^{p}=\delta_{q}^{p}+O\left(r^{\prime \prime}-r^{\prime}\right)$.
On differentiating these two-point tensors with respect to $s \in\left[s^{\prime}, s^{\prime \prime}\right]$, we get due to (4.1)
$\mathrm{DH}_{\cdot \mathrm{p}}^{\mathrm{k}} / \mathrm{ds}=\mathrm{D}^{2} \mathrm{H}_{\cdot \mathrm{p}}^{\mathrm{k}} / \mathrm{ds} \mathrm{s}^{2}=\mathrm{O}\left(\rho^{\prime \prime \prime}-\rho^{\prime}\right)$,
$\mathrm{D} \Lambda_{\cdot \mathrm{q}}^{\mathrm{p}} / \mathrm{ds}=\mathrm{D}^{2} \Lambda_{\cdot \mathrm{q}}^{\mathrm{p}} / \mathrm{ds}{ }^{2}=\mathrm{O}\left(\mathrm{r}^{\prime \prime}-\mathrm{r}^{\prime}\right)$.
By vertue of eq. (3.3) from ${ }^{\prime 3 /}$, sect. II. 3 we have
$\int_{r^{\prime}}^{r^{\prime \prime}} \Lambda_{\cdot q}^{p} \dot{\gamma}_{s}^{q}(r) d r=\dot{\gamma}_{s}^{p}\left(r^{\prime}\right)\left(r^{\prime \prime}-r^{\prime}\right)+O\left(\left(r^{\prime \prime}-r^{\prime}\right)^{2}\right) \quad$ (this result is a direct consequence of (6.1) and (6.6) (see below)); thus from (4.3) and (6.1), we find
$\mathrm{h}_{12}(\mathrm{~s}, \mathrm{x})=\xi-\mathrm{O}\left(\left(\mathrm{r}^{\prime \prime}-\mathrm{r}^{\prime}\right)^{2}\right)$,
where the vector

$$
\begin{equation*}
\xi:=\xi(s):=\dot{\gamma}_{s}\left(r^{\prime}\right)\left(r^{\prime \prime}-r^{\prime}\right) \tag{6.4}
\end{equation*}
$$

is called an infinitesimal (local) deviation vector and evidently has an order of $O\left(r^{\prime \prime}-r^{\circ}\right)$

Now substituting (6.1)-(6.3) into (4.9), we get

$$
\begin{align*}
& \frac{D^{2}}{d s^{2}} \xi^{k}=O\left(\rho^{\prime \prime}-\rho^{\prime}\right)+O\left(\left(r^{\prime \prime}-r^{\prime}\right)^{2}\right)+\int_{r^{\prime}}^{r^{\prime \prime}} \mathrm{drl} \mathrm{O}\left(\mathrm{r}^{\prime}-\mathrm{r}\right)+ \\
& +R_{\cdot i j \ell}^{k} \quad\left(\gamma_{s}(r)\right) \gamma_{s}^{\prime j}(r) y_{s}^{\prime j}(r) \dot{y}_{s}^{\ell}(r)+T_{\cdot j \ell}^{k}\left(\gamma_{s}(r)\right) F_{s}^{j}(r) \dot{\gamma}_{s}^{\ell}(r)+ \\
& \left.+\mathrm{DF}_{\mathrm{s}}^{\mathrm{k}}(\mathrm{r}) / \mathrm{dr}+\gamma_{\mathrm{s}}{ }^{\mathrm{j}}(\mathrm{r}) \mathrm{D}\left(\mathrm{~T}_{\cdot j}^{\mathrm{k}} \ell\left(\gamma_{\mathrm{s}}(\mathrm{r})\right) \dot{\gamma}_{\mathrm{s}}^{\ell}(\mathrm{r})\right) / \mathrm{ds}\right] . \tag{6.5}
\end{align*}
$$

But for any smooth function $f:\left[r^{\prime}, r^{\prime \prime}\right] \rightarrow \mathbb{R}$ the following equality is fulfilled:

$$
\int_{r^{\prime}}^{\prime \prime} f(r) d r=f\left(r^{\prime}\right)\left(r^{\prime \prime}-r^{\prime}\right)+O\left(\left(r^{\prime \prime}-r^{\prime}\right)^{2}\right) .
$$

(Proof: expand the integral in (6.6) into a Taylor series with respect to the difference ( $\mathrm{r}^{\prime \prime}-\mathrm{r}^{\prime}$ ) at the point $\mathrm{r}^{\prime}$ ).

Thus, applying (6.6) to the integral in (6.5) and using (6.4), we find the following final result:

$$
\begin{aligned}
& \frac{D^{2}}{d s^{2}}{ }_{x} \xi^{k}=R_{\cdot i j \ell}^{k}\left(\gamma_{s}\left(r^{\prime}\right)\right) \gamma_{s}^{\prime i}\left(r^{\prime}\right) \gamma_{s}^{\prime j}\left(r^{\prime}\right) \xi^{\ell}+ \\
& +\frac{D F_{s}^{k}(r)}{d r}{ }_{r=r^{\prime}}\left(r^{\prime \prime}-r^{\prime}\right)+T^{k}{ }_{j} f^{f}\left(\gamma_{S}\left(r^{\prime}\right)\right) F_{s}^{j}\left(r^{\prime}\right) \xi^{\ell}+ \\
& +\gamma_{S}^{\prime j}\left(r^{\prime}\right) \frac{\mathrm{D}\left(\mathrm{~T}_{-\mathrm{j}}^{\mathrm{k}}\left(\gamma_{\mathrm{S}}\left(\mathrm{r}^{\prime}\right)\right) \xi^{\ell}\right)}{\mathrm{ds}}+\mathrm{O}\left(\rho^{\prime \prime}-\rho^{\prime}\right)+\mathrm{O}\left(\left(\mathrm{r}^{\prime \prime}-\mathrm{r}^{\prime}\right)^{2}\right) .
\end{aligned}
$$

If we neglect here the terms of an order of $O\left(\rho^{\prime \prime}-\rho^{\prime}\right)$ and $O\left(\left(r^{\prime \prime}-r^{\prime}\right)^{2}\right)$, we shall get the first order deviation equation (or all the same the first order equation of motion) whose solutions (with respect to $\xi$ ) are the first approximation to the solutions of the general equation (4.9).

From (6.7) one can easily get the derived in ${ }^{\prime}$ equation of relative motion (deviation equation) of two point particles which are moving along the curves $x(s, v)$ and $x(s, v+d v)(s$ parametrizes their trajectories) of the family of $\mathrm{C}^{2}$-curves $\mathrm{x}(\sigma, r)$. Defining $\mathrm{u}^{\alpha}:=\partial \mathrm{x}^{\alpha}(\mathrm{s}, \mathrm{v}) / \partial \mathrm{s}$, substituting $\mathrm{x}(\mathrm{s}, \mathrm{v})$ for $\gamma_{s}(r), v$ for $r^{\prime}, v+d v$ for $r^{\prime \prime}, u^{i}$ for $\gamma_{s}^{\prime \prime}(r)$ and putting $\rho^{\prime \prime}=\rho^{\prime}$, we get from (6.7):

$$
\begin{align*}
\frac{D^{2} \xi^{k}}{\partial s^{2}} & =R_{\cdot i, j p}^{k} u^{i} u^{j} \xi^{\gamma} \cdot\left(\frac{D F_{s}^{k}(v)}{\partial v}\right) d v+ \\
& +T_{\cdot j \rho}^{k} F_{s}{ }^{i}(v) \xi^{\beta}+u^{j} \frac{D\left(T^{k} \cdot j \rho \xi^{\ell}\right)}{\partial s}+O\left((d v)^{2}\right) \tag{6.8}
\end{align*}
$$

where

$$
\begin{equation*}
\xi^{\alpha}=\frac{\partial x^{\alpha}(s, v)}{\partial v} d v, \quad F_{s}^{k}(v)=\frac{D u^{k}}{\partial s} \tag{6.9}
\end{equation*}
$$

To get from (6.8) the deviation equation derived in ' $6^{\prime \prime}$, one has to neglect in (6.8) the terms $\mathrm{O}\left((\mathrm{dv})^{2}\right)$ and to put $\mathrm{T}^{\mathrm{k}} \mathrm{ij}^{\prime}=0$ hecause in ${ }^{\prime 64}$ the usual space-time of general relativity without torsion is used.

## 7. CONCLUSION

In this article we have shown that the equations of relative motion of two point particles in the field of external force $F_{s}(r)$, as they are observed from a third point particle (observer) which is influenced by another force field F(r), are a special case of the generalized deviation equation.

One can give many examples, e.g., for the force field $F_{s}(r)$ (cf. ${ }^{\prime \prime \prime}$ ). For instance, if the observed particles are in an electromagnetic field described by a tensor $\mathrm{F}_{\mathrm{ij}}$, we have to put

$$
\begin{align*}
& F_{s}^{j}\left(r^{\prime}\right)=\frac{e^{\prime}}{m^{\prime}} g^{j k}\left(\gamma_{s}\left(r^{\prime}\right)\right) F_{k \ell}\left(\gamma_{s}\left(r^{\prime}\right)\right) \gamma_{s}^{\prime \ell}\left(r^{\prime}\right), \\
& F_{s}^{j}\left(r^{\prime \prime}\right)=\frac{e^{\prime \prime}}{m^{\prime \prime}} g^{j k}\left(\gamma_{s}\left(r^{\prime \prime}\right)\right) F_{k \ell}\left(\gamma_{s}\left(r^{\prime \prime}\right)\right) \gamma_{s}^{\prime \ell}\left(r^{\prime \prime}\right), \tag{7.1}
\end{align*}
$$

where $e^{\prime}, m^{\prime}$ and $e^{\prime \prime}$ and $m^{\prime \prime}$ are the electric charges and the masses of the particles 1 and 2 , respectively, and we have admitted that $L_{n}$ is endowed (may be independent of the connection) also with a metric tensor with contravariant components $\mathrm{g}^{\mathrm{jk}}$.

The values of $\mathrm{F}_{\mathrm{S}}(\mathrm{r})$ for $\mathrm{r} \in\left(\mathrm{r}^{\prime}, \mathrm{r}^{\prime \prime}\right)$ have to be chosen from physical considerations conserning the choice of the curves $\gamma_{s}(r)$, whose physical meaning was mentioned in ${ }^{\prime \prime}$, sect.TV.2.

The author thanks Professor N.A.Chernikov for the useful discussions.

REFEREVCES

1. Alexandrov A.N., Pyragas K.A. - TMF, 1979, Vol.38, No.1, p.71-83 (in Russian).
2. Alexandrov A.N., Pyragas K.A., Pyragas L.E. - Izvestiya VUZ, Sec. Physics, 1983, No.8, p.38-45 (in Russian).
3. Iliev B.Z. - Bulg.J.Phys., 1986, Vol.13, No.6.
4. Iliev B.Z. - Bulg.J. Phys., 1987, Vol.14, No.1.
5. Pyragas L.E. - Izvestiya VUZ, Sec.Physics, 1978, No.12, p. 21 (in Russian).
6. Weber J. - General Relativity and Gravitational Waves, New York, 1961 (ch.8, sect. 1).

НЕТ ЛИ ПРОБЕЛОВ В ВАШЕЙ БИБЛИОТЕКЕ？
Вы мохете получить по почте перечисленные нихе книги，

## если они не были заказаны ранее．

| 29－82－664 | Труды совешания по коллективнм методам усхорения．Дубна， 1982. | 3 p .30 k. |
| :---: | :---: | :---: |
| 43，4－82－704 | Труды IV Международной школы по нейтронной Физике．Дубна， 1982. | 5 p． 00 K． |
| A11－83－511 | Труди совешания по системам и методзм аналитических ешчиспений на ЭВМ и их применению －теоретической Физике．Аубна， 1982. | 2 р． 50 \％． |
| 47－83－644 | Трудн Мехдународмой шкопн－сеяинара по физике тптелих понов．Алуита， 1983. | 6 p． 55 |
| 42，13－83－689 | Трудн рабочего совещанкя по проблемам излучения и детектиродания гравитаиионнмх волн．Дубна， 1983. | 2 D． 00 |
| 413－84－63 | Трудм XI Международного симпозиума по ядерной злектронике．Ератислзпа， Чехословакия， 1983. | 4 p． 50 ～ |
| A2－84－366 | Трудм 7 Мешдународмого совемания по проблемам көантовой теории поля．Апукта， 1984. | 4 P． 30 |
| ベ，さ－ベャーラララ | іруде $\overline{\text { iII }}$ петдународного семинара по проблемам физики писоких знергий．Дубна， 1984. | 5 p． 50 k ． |
| 817－84－850 | Труди Ш1 Мехдународного симпозиума по избранным проблемам статистической механики．Дубна， 1984. $/ 2$ тома／ | 7 ロ． 75 к． |
| A10，11－84－818 | Трудн V Международного совещания по про－ блемам математического моделирования，про－ граммированио м матенатическин нетодам реше－ ния физических задач．Дубна， 1983 | 3 p .50 k. |
|  | Тоуди IX всесоюзного совещания по ускорителям заряженных частиц．Дубна， $1984 / 2$ тома／ | 13 p .50 k. |
| 24－85－851 | Труды Междумародной школы пс структуре ядра，Алушта， 1985. | 3 p． 75 |
| Д11－85－791 | Труды Международного соөещания по аналитическим вынислениям на ЗвМ и их примененио в теоретиие－ ской физике．Дубна，1985． | ． |
| Д13－85－793 | Труды хп Международного симпозиума по ддерной электронике．Дубна 1985. | 4 p． 80 |
| 23，4，17－86－747 | Труды у Международной школы по нейтронной физике．Алушта， 1986. | 4 р． 50 к． |

Илиев Б． 3.
E2－87－78
Уравнение девиации как уравнение движения
На основе понятия о неинфинитезимальном векторе девиа－ ции дан новый вывод в пространствах аффинной связности об－ обденного уравнения девиации в случае неявного задания／по－ средством дифференциальных уравнений второго порядка／участ－ вуюмих в его определении кривых，которые могут быть как геодезические，так и негеодезические．Показано，что част－ ньм случаем этого уравнения являются уравнения относитель－ ного движения двух точечных частиц в пространствах аффин－ ной свяэности в поле внешних сил．В качестве примеров рас－ смотрены＂евклидовый случай＂и общее инфинитезимальное／ло кальное／уравнение девиации．

Рабога выполнена в Лаборатории теоретической физики Оияи．

Сообщение Объединенного института пдерных исследований．Дубна 1987

## Iliev B．Z．

The Deviation Equation as an Equation of Motion
On the basis of a noninfinitesimal deviation vector a new derivation（in spaces with the affine connection）of a ge－ neralized deviation equation is given for the case when the．curves appearing in it，which can be geodesic as well as nongeodesic，are given implicitly（by second order dif－ ferential equations）．It is shown that the equations $q f$ relative motion of two point particles in spaces with the affine connection in the field of external forces are spe－ cial cases of the generalized deviation equation．As exam－ ples the＂Euclidean case＂and the general infinitesimal （local）deviation equation are considered．

The investigation has been performed at the Laboratory of Theoretical Physics，JINR．

Communication of the Joint Institute for Nuclear Research．Dubna 1987

