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# RELATIVISTIC CORRECTIONS TO s AND p WAVES OF PION-NUCLEON <br> SCATTERING <br>  

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## 1. Introducton

It has been shown in refs. 1,2 that to describe correctly the energy dependence of $s$-and $p$-phase shifts of the pion nucleon scattering it is necessary to take into account the pion-pion interaction. However, consideration has been carried out in a static approximation.

The effect of the relativistic corrections on the estimation of the pion-pion interaction in treating pion- nucleon scattering has been recently considered ${ }^{3-7}$ in detail.

In papers ${ }^{3,4}$ it has been shown that the formula CGLN for Red $(\mathbb{S})$ does not describe experimental data even at energies not higher than 100 MeV . The Born term is especially sensitive to relativistic corrections. In refs. ${ }^{5,6}$ the authors have used the Cini- Fubini method. Comparing their results for $s$-wave with the nonrelativistic approximation the authors are led to the conclusion that the non-relativistic approximation can lead to an incorrect extraction of the pion-pion interaction. While for $p$-wave the relativistic corrections become remarkable ( $\approx 30 \%$ ) only at energies 250 MeV .

The problem of pion-nucleon scattering at low energies is reviewed in ref. ${ }^{7}$.

In the present paper we consider the $s^{(-)}$- wave of the pion- nucleon scattering with account of the relativistic corrections and additional inclusion of the $s$-wave into the unitarity condition (in addition to the large $\mathrm{p}_{33}$ wave ). The method presented in ref. ${ }^{1,2}$ has been used. It differs from other approaches ${ }^{3,6}$ : one employs the complex plane $\nu=q^{2}$ instead of $s$ and one find an analytic expression for $R e^{(\cdot)}$ satisfying the crossing symmetry relations instead of detailed calculation of the energy behaviour of phase shifts. The dispersion relations permit to determine the functional dependence of $R \mathrm{Re}_{\mathrm{s}}^{(-)}$on the parameters like scattering length, $\rho$ - meson position, high-energy contribution and so on.

Already in the non-relativistic approximation this method has led to satisfactory agreement with the experimental data on $s$ and $p$ phase shifts of pionnucleon scattering.

As long as the pion-pion contribution to pion-nucleon scattering obeys the symmetry conditions ${ }^{2}$, it is sufficient to get information about the pion-pion contribution to $s$ wave to determine its contributions to other pion-nucleon scattering waves.

The comparison of the relativistic calculation of $s$ - phase shift with the non-relativistic one was made for the same values of the parameters (scattering
length pion-pion resonance position in $J=T=1$ state and pion-nucleon interaction constants).

## 2. Equation for $f_{s}^{(-)}$Partial Wave and Comparison with the Non-Relativistic Calculation

In deriving equation for $f_{s}^{(-)}$-wave use has been made of Eqs. (4.1)-(4.3) from ref. ${ }^{1}$. Expression for $\operatorname{lm} u(\nu, \pm 1)$ and $\operatorname{Im} B^{(-)}(\nu \pm 1)$ were taken in the form:

$$
\begin{align*}
& \frac{\operatorname{lm} a\left(\nu_{ \pm} \pm 1\right)}{4 \pi}=\frac{1}{4[K(\nu)+\nu]}\left[\frac{W+M}{E+M} \ln f_{E}^{(-)}+\left(-3 \frac{W+M}{E+M}+\frac{W-M}{E-M}\right) \operatorname{Im} f_{D A 2}^{(-)}\right]  \tag{2.1}\\
& \frac{\operatorname{In} B^{(-)}\left(\nu_{0} \pm 1\right)}{4 \pi}=\frac{\operatorname{Im} f_{a}^{(\rightarrow)}}{E+M}+\left(\frac{3}{E+M}-\frac{1}{E-M}\right) \operatorname{Im} f_{D B A}^{(\rightarrow)}: \tag{2.2}
\end{align*}
$$

where

$$
a(\nu, \pm 1)=\frac{\Lambda^{(-)}(\nu, \pm 1)}{s-s} ; \quad K(\nu)=\sqrt{\left.(\nu+1)(\nu+M)^{2}\right)} ; \quad W=E+\omega
$$

$E, \omega$ are the energies of nucleon and $\pi$-meson in the c.m.s. respectively, $\nu$ is the squared momentum of $\pi$ meson (or nucleon in the c.ms..), $\pm 1$ are the values of the cosine of the forward or backward scattering angle.

The account of the relativistic corrections and $\operatorname{Im} f_{s}$ in the unitarity cond:tion yields to the following expression for the scattering length

$$
\begin{align*}
& a^{(\rightarrow)}=\frac{2 M}{M+1} \frac{f^{2}}{1-\frac{1}{4 M^{2}}}  \tag{2.3}\\
& \quad+\frac{M}{M+1} \frac{1}{\pi} \int_{0}^{\infty} \frac{d \nu^{\prime}}{\nu^{\prime}}\left[\frac{1}{\omega^{\prime}}+\frac{1}{E^{\prime}}\| \| \mathrm{Im} f_{:}^{(-)}+2 \operatorname{Im} f_{p: / 2}^{(\rightarrow)}\right.
\end{align*}
$$

This equation and analogous relations for the scattering wave lengths $a_{(\rightarrow)}^{(\rightarrow)}$ and $a_{8}^{(\rightarrow)}$ obtained from equations for $P_{(-)}^{(-)} \quad$ and $R=f_{i, 2}^{(-)}$has been used in deriving a final expression for Re $f_{a}^{(-)}$. Using EqE. (2.1)-(2.3) and dispersion relations (4.10), (4.2) from ref. ${ }^{1}$ we get the followi y equation for $\operatorname{Re} f_{a}^{(-)}$:

$$
\begin{equation*}
R e f_{t}^{(\rightarrow)}=\frac{E+M}{4 W} \sum_{i=1}^{s} f_{i} \tag{2.4}
\end{equation*}
$$

where

$$
f_{1}=2 \frac{M+1}{M}(W-M) a^{(-)}
$$

$$
\begin{aligned}
& \mathrm{f}_{2}=\frac{2 \nu}{\pi} \int_{0}^{\infty} \frac{d \nu^{\prime}}{\nu^{\prime}\left(\nu^{\prime}-\nu\right)} \frac{\omega \mathrm{E}^{\prime}}{a^{\prime}} \operatorname{lm} \mathrm{f}_{\mathrm{a}}^{(-)}(\nu) \\
& f_{3}=[F(\nu)-1] \left\lvert\,\left\{2 \frac{M+1}{M}(W-M) a^{(-)}-f^{2} \frac{W-M}{M^{2}-1 / 4}+\frac{1}{\pi} \int_{0} d \nu^{\prime} \cdot \operatorname{Im} f_{D B / 2}^{(-)} G(\nu)\right\}\right. \\
& f_{4}=\frac{1}{\pi} \int_{0}^{\infty} \mathrm{d} \nu^{\prime} \cdot \operatorname{Im} f_{D N a}^{(-)}\left[\frac{\nu^{\prime}}{a^{\prime}\left(\omega+\omega^{\prime}\right)}\left(\frac{2}{\omega^{\prime}}-\frac{1}{E^{\prime}+M}\right)-\frac{\omega\left(a^{\prime}+K\right.}{a^{\prime} K^{\prime}\left(E^{\prime}+E\right)}\left(1+\frac{\omega^{\prime}\left(E^{\prime}+M\right)}{\nu^{\prime}}\right)\right. \\
& +\frac{E^{\prime}+M}{E^{\prime}+E} \frac{1}{\nu^{\prime} a^{\prime}}\left(2 \nu^{\prime}+\omega \frac{\nu^{\prime}-\nu}{E^{\prime}+E}\right)-\frac{4}{a^{\prime}}-\frac{3 \omega^{\prime}}{a^{\prime}\left(E^{\prime}+M\right)}+3 \frac{\omega^{\prime} \psi^{\prime}}{a^{\prime} K^{\prime}\left(E^{\prime}+M\right)}\left(\frac{E^{\prime}}{\omega+\omega^{\prime}}+\frac{\omega}{E^{\prime}+E}\right) \\
& \left.+3 \frac{\omega \nu^{\prime}}{a^{\prime} K^{\prime}\left(E^{\prime}+E\right)}+\frac{A B}{K^{\prime}}+G\left(\nu^{\circ}\right)\right] \\
& f_{0}=f^{2}(W-M)\left[\frac{1}{M^{2} \omega^{2}-1 / 4}\left(1-8 \nu M \frac{a+K+\omega}{4 a^{2}-1}\right)-\frac{1}{M^{2}-1 / 4}\right] \\
& a=\mathrm{K}+\nu \\
& A=1-\left(1-\frac{1}{M^{2}}\right) \frac{\nu M^{2}}{\left(a^{\prime}+a\right)^{2}} \frac{\left(1+\frac{2 a}{M^{2}+1}\right)\left(1+\frac{2 a^{\prime}}{M^{2}+1}\right)}{1+2 \frac{a a^{\prime}+M^{2}}{(a+a)\left(M^{2}+1\right)}} \\
& B=\frac{3}{E^{\prime}+M}\left[\frac{a}{a^{\prime}}\left(W^{\prime}+M\right)+W-M\right]+\frac{E^{\prime}+M}{\nu^{\prime} a^{\prime}}\left[\nu \sigma^{\prime}-\nu^{\prime} \omega-\frac{\omega \omega^{\prime}\left(\nu^{\prime}-\nu\right)}{E^{\prime}+E}\right] . \\
& +\frac{a}{a^{\prime}}-\frac{\nu}{\nu^{\prime}} \frac{E^{\prime}+M}{E+M} \\
& G\left(\nu^{\prime}\right)=6 \frac{K \omega^{\prime}}{K^{\prime} \nu^{\prime}\left(E^{\prime}+M\right)}+6 \frac{W-M}{\nu^{\prime}\left(E^{\prime}+M\right)}-4 \frac{W-M}{\nu^{\prime} E^{\prime} \cdot}-5 \frac{\omega\left(\nu^{\prime}-\nu\right)}{\nu^{\prime} \cdot K^{\prime} \cdot\left(E^{\prime}+E\right)} \\
& +3 \frac{W-H}{R^{\prime}\left(E^{\prime}+M\right)}-\frac{\nu}{\nu^{\prime}} \frac{\left(E^{\prime}+M\right)}{K^{\prime}(E+M)}-\frac{4 \nu}{\nu^{\prime} \omega^{\prime}(E+M)}+\frac{\nu \omega}{\nu^{\prime} \cdot K^{\prime}(E+M)}
\end{aligned}
$$

in which the sense of each term of (2.4) is clear from the definition.

In the limiting case $M \rightarrow \infty$ the functions $f_{i}$ are equal to

$$
\begin{align*}
& \mathrm{f}_{1}=2 \omega a^{(-)} \\
& \mathrm{f}_{2}=\frac{2 \nu}{\pi} \int_{0}^{\infty} \frac{\mathrm{d} \nu^{\prime}}{\nu^{\prime} \cdot\left(\nu^{\prime}-\nu\right)} \frac{\omega}{\omega^{\prime} \cdot} \operatorname{Im} \mathrm{f}_{s}^{(\rightarrow)} \\
& \mathrm{f}_{3}=[\mathrm{F}(\nu)-1] 2 \omega \cdot \mathrm{a}^{(-)}  \tag{2.5}\\
& \mathrm{f}_{4}=\mathrm{f}_{5}=0
\end{align*}
$$

If,, in addition, we neglect the contribution from $\operatorname{Im} f(-)$ vistic approximation for $\operatorname{Re} \mathrm{f}_{\mathrm{s}}^{(-)} \quad\left(\right.$ see ref. ${ }^{1}$ )

$$
\begin{equation*}
\operatorname{Ref} \mathrm{f}_{\mathrm{a}}^{(\rightarrow)}=\omega \mathrm{a}^{(\rightarrow)} \mathrm{F}(\nu) \tag{2.6}
\end{equation*}
$$

Table $1 \underset{(-)}{\text { gives }}$ the values of each contribution of $f_{i}$ separately to the expression for $R e f s$. The comparison with the non-relativistic expression (2.6) is also presented. Table 1 shows that the relativistic corrections are small in the considn dered region of momenta $0 \leq \nu \leq 3$. The relativistic corrections are largest in the resonance region $\eta \approx 1.5$. They change as desired the non-relativistic curve and make it closer to the experimental data (see Fig.1).

Table 1 representes the behaviour of the relativistic corrections to $R e f_{s}^{(-)}$. The function $f_{2}$, associated with the account of $\operatorname{Im} f_{s}(-)$ in $\operatorname{lmA}(-)$ and $\operatorname{Im} B^{(-)}$is small. Of importance is the compensation of functions $f_{4}$ and $f_{5}$ which leads to $\mathrm{Re} \mathrm{f}_{\mathrm{s}}^{(-)}$being mainly determined by $\mathrm{f}_{1}$ and $\mathrm{f}_{3}$ i.e.

$$
\begin{equation*}
\operatorname{Ref}_{\mathrm{s}}^{(-)}=\frac{E+M}{2 M} \frac{M+1}{W}(W-M) a^{(-)} F(\nu) \tag{2.7}
\end{equation*}
$$

Thus, the account of the relativistic corrections does not change practically the way in which the function depends $\operatorname{Ref}_{s}^{(-)}$on the parameters $a_{a}$ and $t_{r}$ and results in the appearance of some kinematic fadtor (see Table II).

Table III/gives the comparison of relativistic and non-relativistic pion-pion contributions to $f_{s}^{(-)}$-wave. The account of the relativistic corrections does not affect essentially the pion-pion interaction.

## Conclusion

1. The calculation shows that in the used approach ${ }^{1,2}$ the relativistic corrections are small through the whole energy region under consideration.
2. The account of the pion-pion interaction is essential to describe correctly the energy behaviour of $s^{(-)}$and $p_{1 / 2}^{(-)}$phase shifts of pion- nucleon scattering.
3. In more exact experimental determination of the pion- nucleon scattering phase shifts at energies up $300-400 \mathrm{MeV}$ there is a reliable possibility to single out the pion-pion interaction and to determine its parameters (scattering wave length and resonance position).
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Fig. 1

Table_L

| $\mathrm{Fef}_{\mathrm{s}}$ | $\eta=\sqrt{\nu}$ | $\mathrm{f}_{1}$ | $f_{2}$ | $f:$ | 1. | 13 | $\eta$ Re $f_{s}^{(\text {(4) }}$ reloty | $\eta \operatorname{Re} \int_{i-j}^{\operatorname{To}} \text { (felaty) }$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| ) | 0 | 0,184 | 0 | 0 | 0 | 0 | 0 | I |
| 1,043 | 0,5 | 0,209 | 0,00049 | -0,0109 | 0,019 | -0,0114 | 0,0404 | 0,94 |
| 0,096 | I | 0,274 | 0,00012 | -0,0469 | 0,007 | -0,0271 | 0,085I | 0,89 |
| 0,206 | 2 | 0,465 | 0,0097 | -0.193 | 0,025 | -0,0407 | 0,190 | 0,92 |
| 0,288 | 3 | 0,699 | 0,0264 | -0,409 | 0,040 | -0,044I | 0,287 | I |

Table II

| $\eta$ | $\begin{aligned} & \operatorname{Refe}_{6}^{(i)} \\ & (2.6)^{2} \end{aligned}$ | $\begin{aligned} & \text { Ref. } \\ & (27) \end{aligned}$ | $\begin{aligned} & \text { (26) } \\ & \text { (27) } \end{aligned}$ |
| :---: | :---: | :---: | :---: |
| 0 | 0,08 | 0,08 | I,00 |
| 0,5 | 0,086 | 0,086 | I,00 |
| I | 0,096 | 0,096 | I, ט0 |
| 2 | 0,103 | 0,100 | I,03 |
| 3 | 0,096 | 0,89 | I,08 |

Table III

| $\eta$ |  |  | $\begin{aligned} & G_{m \pi} \text { (relatuv.) } \\ & G_{g a} \text { (noirelativ.) } \end{aligned}$ |
| :---: | :---: | :---: | :---: |
| 0 | 0 | 0 | I |
| 0,5 | -0,00384 | -0,00467 | I, 21 |
| I | -0,0174 | -0,0193 | I,II |
| 2 | -0,0753 | -0,0718 | 0,952 |
| 3 | -0,157 | -0,137 | 0,873 |

