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ON PHOTON-NUCLEON SCATTERING

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## Abstract

An analysls of elastic $\quad y=$ ray scattering by protons at an energy up to 300 MeV has bean made by using the disparsion relations (d. r.) Six dispersian relations were made use of to estimate the real parts of the ariplitudes at $Q^{2}=0$. The pion phatoproduction has been taken into account in a wider than eariler onetgy region, It has been proved that the subtraction is necessary in the dispersion relations (at any rate for one amplitude). Five subtraction constants are determined by the low energy limit and expressed in terms of the charge and the nucfeon matinetic monent. The differential cross sections and the recoli nucleon polarization have been estimated.
$Y-N$ scaftering at high energies is being discussed.

Following Gell-Mann, Goldberger and Thirting/1/ the dispersion relations (d.r.) for $\gamma-\mathrm{N}$ - acattering the validity of which in the $e^{2}$ - approximation was proved strictly by Bogolubov and Shirkov $/ 2 /$ were applied to analysis of the experimental data in a number of papers $/ 3-7 /$. Cini and Strofolini $/ 3$ / were the first to make the calculations of the cross sections for the forward scattering at the energies of $\boldsymbol{\gamma}$-rays up to 210 VeV . Some qualitative features of the energy dependence of the cross section for ford acattering were pointed out earlier in $/ 1 /$, as well es in ${ }^{/ 8 /}$.

Capps $/ 4 /$ considered $\quad \gamma-N$ scattering at arbitrary angles with account of the minimutm number of angular momentum states. In doing this, he made use of some unpublished results of Gell-Mann and J. Mathews.

Akiba and Sato $/ 5 /$ treated the non-zero angle scattering. To calculate the subtraction constants in some dispersion relations the perturbation theory was employed.

In Ref. ${ }^{/ 6 /}$ the dispersion relations were considered in detail for all six invariant functions characterizing the amplitude for $y-N$ scattering, and the dispersion analysis was made in the energy region up to 200 MeV , some recoil effecte being neglected. It has been shown that the account of the photoproduction of pions in the s-state leads to noticeable changes in the threshold region. At the same time the agreement between the dispersion analyais and experimental data improves. Casp dependence arised in the amplitudes and cross sections near the threshold for pion photoproduction.

In opite of some differences in the published papers which are due to the different assumptions on the number of subtractions in the d.r. and on the maximum moment of the considered states, the common feature for all of them was the impossibility of obtaining good agreement with the experimental data in the energy range of about $160-200 \mathrm{MeV}$.

In a number of papers $/ 9,7,10$ / an attempt has been undertaken to do away with the discrepancy by taking into account the Low diagram/11/. However, a direct measurement of the $\pi^{\circ}$-meson lifetime ${ }^{/ 12 / \text { and }}$ the analysia of the problem concerning the sign of the pole amplitude and its magoitude ${ }^{/ 13 /}$ has led to the
fact that the account of the Low amplitude cannot affect essentially the result of the analysis.
In view of the differences between the results of the analysis and the available experimental data this paper deals with the analysis of $\gamma \cdot \mathrm{N}$-acattering based on disperaion relations. In the course of this analysis, hesides the account of the pion photoproduction in the s-state, we are considering the contribution of the high energy region more thoroughly. Then we are treating the problem on the number of subtractions in the dispersion relations and by having taken the nucleon recoil into account, we are estimating the quantities. $R_{i}(\nu)$ for $Q^{2}=0$ introduced earlier.

## 2

The connection of the invariant functions $\quad T i\left(\nu, Q^{2}\right)$ with the amplitudes $R i\left(\nu, Q^{2}\right)$ in the center-of-mass system (c.m.s.) is given in formula (1) of paper $/ 14 /$ (which is further referred to as paper A). The definitions of $T i\left(\nu, Q^{3}\right)$ and $R i\left(\nu, Q^{2}\right)$ see in $/ 13 /$ (which is further referred to as paper B ). The notations adopted in the present paper coincide with those of the papers A and B . Here the amplitudes in the c.m.s. are designated by $R_{i}$ without additional asterisks.

As far as by the optical theorem

$$
\begin{equation*}
I m\left(R_{1}+R_{g}\right)=\frac{v_{2} g_{1}}{4}=\frac{V^{2}-M^{2}}{2 I} \quad \frac{a_{1}}{4}, \tag{1}
\end{equation*}
$$

then under the assumption

$$
a(w) \rightarrow \text { Conae } \quad \text { by } \quad W \rightarrow \infty
$$

asymptotically by $\qquad$

$$
\begin{equation*}
R_{1}+R_{2} \rightarrow W \mathbb{S}^{\prime} V . \tag{2}
\end{equation*}
$$

Assuming further that by $W \rightarrow \infty$ all $R_{i} \sim \nu$ we get from (1.A), that by $W \rightarrow \infty$

$$
\begin{array}{lll}
T_{y}-T_{s} \rightarrow W^{2} ; & T_{s}+T_{3} \rightarrow W^{2} ; & T_{s} \rightarrow W^{2} \\
T_{y} \rightarrow T_{4} \rightarrow W ; & T_{3}+T_{4} \rightarrow \text { Conat } ; & T_{4}+W
\end{array}
$$

Thus, under the adopted assumptions the diepersion relations for $T_{1}, T_{3}$ and $T_{5}$ must contain one subtraction, while for the quantities $\mathrm{T}_{2}, \mathrm{~T}_{4}$ and $\mathrm{T}_{6}$ the disperaion relations can be written without subtractions.

To estimate the amplitudes $R_{1}+R_{2}, R_{3}, R_{4}, R_{5}+R_{6}$
it is aufficient to write down the dispersion relation for $T_{i}$ at $Q^{2}=0$. When $Q^{2}=0$, the invariant $v$

$$
v=v_{L}-Q^{3} h
$$

is known to pass into the energy of $\quad \gamma$-rays in the labsystem $\nu_{L}$ (which is further designated by $\nu$ ). As is seen from (1.A), for the forward scattering the function $T_{5}$ and $T_{2}+T_{4}$ are reduced to $\Pi_{4}-R_{3}$, so that the dispersion relations for $T_{5}$ and $T_{2}+T_{4}$ at $Q^{2}=0$ are equivalent.

Let us consider the functions

$$
\begin{align*}
& F_{1}\left(\nu_{j}\right)=\frac{1}{2}\left[T_{i}-T_{3}-\nu_{0}\left(T_{2}-T_{4}\right)\right]=\frac{\mathbb{K}_{2}}{N}\left(R_{p}+R_{2}\right) \\
& F_{2}\left(V_{d}\right)=V T_{4}=\frac{V_{2}}{H}\left[R_{3}+R_{4}+2 R_{3}+2 R_{5}\right]  \tag{4}\\
& F_{g}\left(V_{0}\right)=\frac{V_{0}}{2 H}\left(T_{s}+T_{s}\right)=\left(\frac{T_{R}}{M}\right)^{2}\left(R_{3}-R_{8}\right) \\
& F_{4}\left(\nu_{0}\right)=\frac{l}{2}\left(T_{t}-T_{s}\right)=\frac{W^{2}}{W \nu_{0}}\left(R_{y}+R_{4}\right)-\frac{2 W_{0}}{H+W_{*}}\left(R_{t}+R_{y}\right) .
\end{align*}
$$

It is clear from the discussion given above that the dispersion relations for the functions $\quad F_{1+\ldots}, F_{4}$ must contain one subtraction. All the quantities in the righthand side of (4) are in the c.m.s. If we take into account that (at $\mathrm{Q}^{2}=0$ ) the amplitudes in the lab.system ( ${ }^{\prime} \Lambda$ ') are connected with the corresponding quantities in the c.m.s. by

$$
\begin{gather*}
\left(R_{1}+R_{2}\right)^{\Lambda}=\frac{W e}{M}\left(R_{y}+R_{2}\right) \\
{\left[R_{y}+R_{4}+2 R_{y}+2 R_{4}\right]^{\Lambda}=\frac{W_{8}}{N}\left[R_{y}+R_{4}+2 R_{8}+2 R_{6}\right]}  \tag{5}\\
\left(R_{4}-R_{y}\right)^{\Lambda}=\left(\frac{W \rho}{W}\right)^{2}\left(R_{4}-R_{y}\right),
\end{gather*}
$$

then from the dispersion relations for $\mathrm{F}_{1} \ldots \mathrm{~F}_{4}$ we get

$$
\begin{align*}
& D_{1,4}^{A}\left(\nu_{0}\right)-D_{1,4}(0)=\frac{\lambda_{1}^{2}}{\pi} \int_{\nu_{1}}^{\infty} \frac{d v}{\nu^{2}-\nu_{0}^{2}} \frac{A_{t_{0} 4}(\nu)}{\nu}  \tag{6}\\
& D_{2,8}^{A}\left(\nu_{0}\right)-\nu_{0} D_{2, j}^{\prime}(0)=\frac{\lambda_{\nu_{0}}^{s}}{\pi} \int_{L_{1}}^{\infty} \frac{A_{p, 1}(\nu) d \nu}{\nu^{2}\left(\nu^{2}-\nu_{0}^{2}\right)}
\end{align*}
$$

where

$$
\begin{aligned}
& \dot{D}_{i}^{A}=R e\left(R_{1}+R_{2}\right)^{A} \quad D_{1}(0)=-2 / v \\
& D_{2}^{A}=R_{e}\left[R_{3}+R_{4}+2 R_{z}+2 R_{6}\right]^{\wedge} \quad D_{3}^{\prime}(0)=-2 \mu_{z}^{2} \\
& D_{3}^{\Lambda}=R_{e}\left(R_{4}-R_{3}\right)^{\Lambda} \quad D_{3}^{\prime}(0)=-2\left[\mu^{\left.2-\left(\frac{0}{2 I I}\right)^{2}\right]}\right. \\
& D_{4}^{\lambda}=R e F_{4}(v) \quad D_{4}(0)=-\frac{\theta^{2}}{2 \lambda} \lambda(2+\lambda),
\end{aligned}
$$

and the imaginary parts of the corresponding amplitudes are designated by $\mathrm{A}_{\mathrm{i}}\left(\nu_{0}\right)$, the magnetic moment by $\mu=e\left(\frac{1}{+}+\lambda\right.$ and the anonalous magnetic moment of a nucleon by $\mu_{a}$.

If the elements of the amplitude of the pion photoproduction in the states with $1 \leqslant 3 / 2$ are denoted in the c.m.s, by

$$
\begin{array}{lll}
E_{1}-/ e l .1 / 2^{-} / ; & M_{1}-/ n a g n .1 / 2^{+} / ; & E_{2}-/ o l .3 / 2^{+} / ; \\
E_{3}-/ e l .3 / 2^{-} / i & M_{3}-/ \text { nagn. } 3 / 2^{+} / ; & M_{2}-1 \text { magn. } 3 / 2^{-} /,
\end{array}
$$

then the unitarity relations will lead to the equalities

$$
\begin{gather*}
\left.\operatorname{lm} R_{t}=\left.\nu_{0}| | E_{2}\right|^{2}+2\left|E_{3}\right|^{2}+\frac{1}{3}\left|E_{2}\right|^{2} \cos \theta-\frac{1}{6}\left|U_{3}\right|^{2}\right\} \\
\operatorname{lm} R_{3}=\nu_{0}\left\{\left|E_{i}\right|^{2}+\frac{1}{3}\left|E_{2}\right|^{2} \cos \theta-\left|E_{3}\right|^{2}+\frac{1}{12}\left|H_{2}\right|^{2}+R e\left(E_{3}^{*} H_{2}\right\}\right.  \tag{7}\\
\left.\ln R_{s}=-\left.\nu_{0}\left|\frac{1}{6}\right| E_{2}\right|^{2}+\operatorname{Re}\left(E_{2}^{*} H_{y}\right)\right\}
\end{gather*}
$$

which are the generalization of the corresponding equalities in $/ \sigma /$. The expressions for $I_{m} R_{2}$ differ from $I m R_{1}$ by the replacement $\quad E_{i} \not \|_{i}$. Similarly, the expression for $I m R_{4}$ can be obtained from $\mathrm{ImR}_{3}$, and $\mathrm{ImR}_{6}$ from $\mathrm{ImR}_{5}$.

In (7), the total contribution of the photoproduction of $\pi^{+}$and $\pi^{\circ}$-mesons is meant by the module of the amplitude in the right-hand side. Let us note, that if the pion and nucleon mass differences are neglected, the interference terms, e.g., in $E_{0}^{\prime}+\left|E_{4}\right|^{2}$ are vanished due to the isotopic symmetry of the pion photoproduction matrix elements.

At the same time

$$
\begin{align*}
& \left.A_{i}(\nu)=\frac{\nu a_{T}}{h_{\pi}}=\left.\nu| | E_{1}\right|^{2}+\left|U_{1}\right|^{2}+2\left|E_{3}\right|^{2}+\left.2\left|N_{g}\right|^{2} \frac{1}{6}\left|M H_{2}{ }^{2}+\frac{1}{6}\right| E_{2}\right|^{2}\right\} \\
& \left.A_{f}(\nu)=\left.\nu| | E_{t}\right|^{2}+\left|M_{z}\right|^{2}+\frac{1}{3}\left|E_{2}\right|^{2}+\frac{1}{3}\left|\cdot U_{2}\right|^{2}-\left|U_{3}+\frac{1}{2} E_{2}\right|^{2}-\left\lvert\, E_{3}+\frac{1}{2} U_{z}{ }_{2}{ }^{2}\right.\right\} \\
& A_{3}(\nu)=-\nu \frac{W}{V}\left\{\left|E_{2}\right|^{2}-\left|V_{1}\right|^{2}+\left|U_{3}-\frac{1}{2} E_{2}\right|^{2}-\left|E_{3}-\frac{1}{2}, H_{2}\right|^{2}\right\}  \tag{8}\\
& A_{4}(\nu)+\frac{W-i J_{\sigma_{t}}}{t \pi}=\left.W| | E_{d}\right|^{2}+\left|H_{4}\right|^{2}+\frac{2}{3}\left|M_{2}\right|^{2}+\frac{2}{3}\left|E_{2}\right|^{2}-\left|E_{3}-\frac{1}{2} H_{2}\right|^{2}-\left|M_{3}-\frac{1}{2} E_{2}\right|^{2}
\end{align*}
$$

The necessity of considering the dispersion relations with the subtraction for the amplitade $\quad \mathbb{R}_{1}+\mathbf{R}_{2}$ can be proved.

Let us prove that the low energy limit of the amplitude $R_{1}+R_{2}$ contradicts the value of this quantity oblained with the aid of dispersion relations withoat the subtraction.

As is well-known, for $\nu \rightarrow 0\left(Q^{2}=0\right)$

$$
\begin{equation*}
R_{2}+R_{2} \rightarrow-e^{2 / K}<0 \tag{9}
\end{equation*}
$$

the dispersion relations without the subtraction for $R_{1}+R_{2}$ are of the form

$$
\begin{equation*}
\operatorname{Re}\left[R_{1}\left(\nu_{0}\right)+R_{f}\left(v_{0}\right)\right]=\frac{1}{-\pi_{2}^{2}} \int_{\nu_{1}}^{\infty} \frac{\nu^{2} d \nu}{\nu^{2}-v_{d}^{2}} \sigma_{1}(\nu) \tag{10}
\end{equation*}
$$

For $v_{0} \rightarrow 0$, it follow from ( 10 ), that

$$
R e\left(R_{g}+R_{2}\right) \rightarrow+\frac{1}{2 \pi^{2}} \int_{\nu_{1}}^{\infty} \sigma_{f}(v) d v>0
$$

what contradicts equality (9)*.
Thus, the dispersion relation without the subtraction for the emplitude $\mathrm{H}_{1}+\mathrm{R}_{2}$ contradict the relativistic and gauge invariance requiremente on which the low energy limit is based.

Let as note, that possible sum rules connected with the square of the magnetic moment, if the dispersion relations without subtraction are assumed for $F_{2}(v)$, are not in a direct contradietion with the lowenergy limit. As is seen from (6) and (8), the contribution of the resonance state proportional to $\left|\mathrm{M}_{3}\right|^{\mathbf{2}}$ is especially essential here. The same result holds if one takes into account (numerically important) contribution of the photoproduction in the s-state which decreases the effective contribution $\left|\|_{3}\right|^{2}$.

The sum rule for the square of the magnetic moment is very sensitive to the relationship between the photoproduction amplitudes $E_{2}$ and $M_{3}$. Forsome relations (for instance, with $E_{2}=H_{3} / 5 /$ ) one can arrive to a contradiction. However, the information on the analysis of the photoproduction is not so accurate at present to state that the experimental data contradict the sum rule. A further specification of the photoproduction analysis is extremely desirable for obtaining the data on the amplitudes $\quad \mathrm{E}_{2}, \mathrm{E}_{3}$ and $\quad \mathbb{H}_{2}$.

The fact that the dispersion relations without subtractions lead to definite sum rules can be of great interest for ame processes. So, for $\pi \pi$ scattering similar considerations (if applied to the disper-

[^0]sion relations at $2^{2}=0$ ) lead one to the conclusion that the scattering lengths in the s-state at low onergies $a_{0}$ and $a_{2}$ are positive. The same holds for $\pi-K$ and $K-K$ acattering.

## 5

If we consider the properties of the functions*

$$
\begin{gather*}
F_{s}\left(V_{0}\right)=\left(T_{2}-T_{4}\right)^{\prime}  \tag{11}\\
F_{6}\left(V_{4}\right)=\left(T_{2}+T_{4}\right)^{\prime}  \tag{12}\\
F_{7}\left(V_{0}\right)=T_{4}^{\prime}, \tag{13}
\end{gather*}
$$

additionally to those introduced earlier, then one can become sure that $F_{5,6}(\nu)$ are the odd functions of $\nu$, which have no poles, and $F_{7}(\nu)$ is the even function with the second-order pole. At $\nu \rightarrow \infty$.

$$
F_{f, 4,} \rightarrow \nu^{-1 / 2},
$$

so that the dispersion relations for these functions do not contain the subtraction. The use of these dispersion relations may prove useful since in taking into account the photoproduction in the states with $\mathrm{J} \leqslant 3 / 2$ the angular dependence of the amplitudes $\operatorname{Ri}\left(v, Q^{2}\right)$ in the c.m.a. is repreaented as

$$
\begin{aligned}
& R_{y}=\varepsilon_{1}+2 \sigma_{3}+2 \sigma_{2} \cos \theta \cdots{ }_{2}
\end{aligned}
$$

$$
\begin{align*}
& R_{3}=E_{5}-\sigma_{3}+2 E_{2} \cos \theta+\frac{1}{2} T_{3}+C\left(E_{3} W_{3}\right) \\
& n_{4}=\pi_{2}-\pi_{3}+2 \pi_{2} \cos \theta+\frac{1}{2} \xi_{2}+C\left(\pi_{3} \xi_{2}\right)  \tag{14}\\
& \left.R_{s}=-\xi_{2}-C()_{3}\right) \\
& R_{4}=-\pi_{2}-C\left(\xi_{3} n_{2}\right)
\end{align*}
$$

and characterized by the eight function of energy $\delta_{i}$, $M_{i}, C\left(\delta_{3}, F_{2}\right) \quad C\left(\pi_{3} \xi_{2}\right)$, which are expressed in terms of $\mathrm{Ri}^{(\nu, 0)}$ and $\mathrm{R}_{\mathrm{i}}(\nu, 0)$.

It follows from ( 14 ), that if we restrict ourselves to the contribution of the states with $J \leqslant 3 / 2$

[^1]\[

$$
\begin{aligned}
& R_{2}^{\prime}=R_{3}^{\prime}=2 \xi_{3}\left(-\frac{\partial \cos \theta}{\partial Q^{2}} Q_{\left.Q^{2}=r\right)}=-\mathcal{F} \xi_{2} \frac{H_{0}^{2}}{\|^{2} \nu_{o}^{\prime}}\right. \\
& R_{z}^{\prime}=R_{4}^{J}=-1 \eta_{2}^{n} \frac{U_{0}^{2}}{U_{0}^{2} v_{0}^{2}} \\
& R_{s}^{*}=R_{d}^{\prime}=0
\end{aligned}
$$
\]

and

$$
\begin{equation*}
\left(R_{2}+R_{y}\right)^{\prime}=\left(R_{3}+R_{4}\right)^{\prime}-\left(R_{3}+R_{4}+2 R_{3}+2 R_{6}\right)^{\prime} \tag{15}
\end{equation*}
$$

In the low energy limit $^{\prime} 14$

$$
\begin{gather*}
\left(R_{1}+R_{2}\right)^{\prime}=-2 \frac{e^{2}}{W} v^{1}+Q(t) \\
\left(R_{3}+R_{4}\right)^{\prime}--\frac{\theta^{2}}{2 U^{3}}\left[3+2(1+\lambda)^{2}\right]+O(v) \\
\left(R_{3}+R_{1}+2 R_{s}+2 R_{d}\right)^{\prime}--\frac{\theta^{2}}{2 U^{2}}\left(2 \lambda^{2}-2 \lambda-1\right)+O(v) \tag{16}
\end{gather*}
$$

The contradiction of (15) with the low-energy limit (16) implies that the restriction to the stateg with $\mathrm{J} \leqslant 3 / 2$ is not correct even in the low-energy region. The crossing-symmetry conditions introduce the kinematic corrections of the onder of $1 / 4$, what corresponds to the account of the states with higher values of J .

In order to make the analyais with such a high accuracy it is necessary to introduce new functions of energy and to consider a great number of dispersion relations. The infroduction of the Low diagram fails to solve the contradiction mentioned above.

All the estimates of the amplitudes are given here, $\mathrm{I}^{4}{ }^{+}(\nu, 0)$ being neglected.

The results of the calculations for the amplitudes at $Q^{2}=0$ are plotted in the graphs. The energy of $\gamma$ - rays is expressed in the part of the threshold energy $\quad \nu_{t}=150 \mathrm{MeV}$, the values of the amplitu* des in the purt of $e^{2 / M c^{2}}$.

To calculate the differential cross action for the forward scattering

$$
\sigma\left(09=\left|R_{t}+R_{2}\right|^{2}+\left|R_{3}+R_{4}+2 R_{5}+2 R_{4}\right|^{2}\right.
$$

as ie well-known, it suffices to have the amplitudes $H_{1}+\mathrm{H}_{2}$ and $\mathrm{H}_{3}+\mathrm{H}_{4}+2 \mathrm{R}_{5}+2 \mathrm{H}_{6}$.

To estimate $D_{1}\left(v_{0}\right)$ the data an the total cross for $y$-ray interaction with protons have been used These datd menrporated also the second maximum and the rross section tor production of pion pairs. The
 higher than 500 ifc V . The result of the eatimation of the amplitude $H_{1}+H_{2}$ is given in Fig. 2. The main difference with the previous results takes place in the range $1<\nu{ }_{o}<2$, where the values of $D_{1}\left(\nu_{n}\right)$ decrease considetably due to the compensation of the low-energy limit and of the dispersion contribution. Viste. that it is this energ region which turns at to be sensitive to the change of $\mathcal{A}_{1}\left(v_{0}\right)$. The second maximum in $A_{1}\left({ }_{0}\right)$ corresponds to the second maximurn in the photoproduction cross section.
when estimating the real parts of the amplitudes different from $\mathrm{H}_{1}+\mathrm{fl}_{2}$, which require more detailed experimental data on the photoproduction we restrict ourselvea to the energy region up to 300 lifel. As for the amplitude $R_{1}+R_{2}$. it sas finund posable to go further. although with energy the uncertainly of the contribution of the nhotoproduction of pairs (and of larger number) of phons is demonstrated noticeably.

In some papers 15,16 the process of $y-p$ scattering at $300-800 \mathrm{Ve}$ was treated as a diffraction me with $\mathrm{R}_{\mathrm{l}} \mathbb{R}_{i}<\operatorname{Im}{R_{i}}$. An experimental inveatigation of $\gamma-p$ scattering in the range of the second maximum is of interest as a sen tive method for studying the maximum itself.

If, by neglecting all he $\Pi_{i}$, we restrirt ourselves to the inkginary parts of the amplitudes and consider $\mathrm{wnl}_{4}$ the contribution proportional to $\left|\mathrm{F}_{3}\right|^{2}$. then from (:) we obtain immediately that
and the differential crossaection ${ }^{6}$ is equal to

$$
\begin{equation*}
s(1)=n_{g}^{2}\left(7+3 \cos ^{2} \theta\right)=-\frac{\pi_{1}^{2}}{2}\left(7+3 \cos ^{2} \theta\right), \tag{17}
\end{equation*}
$$

what coincides with the \inami's results. The same result holds for the angular distribution, if in (7) only $H_{3}\left(R_{1} \rightarrow R_{2}, R_{3} \rightarrow R_{4}\right)$ is different from zero. For $E_{3}$ and $H_{3}$. simultaneously different from zern (at $\operatorname{Re} \Pi_{i}=0$ )

$$
\begin{equation*}
\sigma(\theta)=\left(R_{3}^{2}+n_{4}^{2}\right) \ldots+\frac{3}{\lambda} \cos ^{2} \theta+10 R_{3} n_{4} \cos \theta . \tag{18}
\end{equation*}
$$

However. as our estimates show the values fle ( $\mathrm{H}_{1}+\mathrm{H}_{2}$ ) cannot be neglected in the region of the second naximum. F'rom this point of view, the second maximum in much different from (3.3) resonance in whore region $\operatorname{Re}\left(I_{1}+n_{2}\right) \ll \operatorname{lm}\left(n_{1}+n_{2}\right)$.

The regults of the calculations for $R_{3}+R_{3}, R_{3}+R_{1}+2 R_{5}+2 R_{6}$, and $n_{5}+R_{6}$ are presented

different from zero. At the same time, for $\left|E_{1}\right|^{2}$ and $\left|M_{3}\right|^{2}$ the energy dependence given in $/ 6 /$ was used, and $\left|\mathrm{E}_{3}\right|^{2}$ was taken to be different from zero in the energy range of $3.1<\nu_{0}<\quad 5.8$.

Note, that even if the imaginary part of $A_{5}+A_{6}$ is absont, the real part of this quantity is different from the low energy limit, as far as the invariant functions $T_{i}(\nu)$ satisfy the dispersion relations.

The values of $\sigma\left(0^{\circ}\right)$ are plotted in Fig. 5 where the result of Cini and Stroffolini for $a_{1}\left(0^{\circ}\right)$ in the c.m.s. is given for comparison. As appreciable difference is observed in the threshold region,

## 7

To entimate $R_{1}-\mathrm{I}_{2}$ and $\mathrm{R}_{5}-\mathrm{I}_{6}$ disporsion relations (6) are not sufficient. Let ue consider the function

$$
\begin{equation*}
P(\nu)=\frac{1}{T^{2}} \phi(\nu)=\frac{\mu^{2} \omega}{2 V^{2}}\left[\left(T_{4}+T\right)^{\prime}-\nu\left(T_{4}+T_{4}\right)^{\prime}\right] \tag{‘19}
\end{equation*}
$$

As one can see from (B.4)

$$
\begin{equation*}
F(\nu)=\frac{Y}{M} 1 R_{z}-R_{z}-\frac{2 M v}{W(W+M)}\left(R,-R_{\ell}\right)! \tag{20}
\end{equation*}
$$

The consideration of the diepersion relations for $F(\nu)$ at the known $n_{3}-R_{4}$ enables us to estimate $\mathrm{A}_{1}-\mathrm{R}_{2}$. In the energy region we are concidering the cocfficient bofore $\left(\mathrm{A}_{3}-\mathrm{A}_{4}\right)$ in $(20)$ is of the order of $v / M$. However, as far as the value $n_{3}-\boldsymbol{R}_{4}$ is great (if compared with $R_{1}-\boldsymbol{H}_{2}$ ), the second component in (20) cannot be neglected.

The function $\phi(\nu)$ introdaced in (19) is an amalytical function of $\nu$ with the cuts along $\nu_{t}<\nu<\infty \quad$ which satisfies the following property of the crossing-symmetry

$$
\begin{equation*}
\phi\{\nu\}=\phi *(\nu \nu) . \tag{21}
\end{equation*}
$$

Thus, for $\nu \ll \nu_{\mathrm{f}}, \phi(\nu)$ is a real function and

$$
\begin{equation*}
\phi(\nu)=a+b r^{\circ}, \tag{22}
\end{equation*}
$$

whereas

$$
\begin{equation*}
F(\nu)=\frac{d(v)}{V^{2}+2 V \nu}=\frac{a}{W^{3}}\left(1-\frac{2 v}{M}\right)+b v^{2}+\ldots \tag{23}
\end{equation*}
$$

We see that the linear term in $\mathrm{F}(\nu)$ is wholly determined by the first tern in (22). Ther, it follows from (20), that for small

$$
R_{1}-R_{2}=-\frac{\theta^{2}}{W}\left(1-\frac{3 v}{W}\right)+O\left(\nu^{2}\right)
$$

and the linear tem in $?_{1}-\mathrm{R}_{2}$ and in $\mathrm{F}(\nu)$ is fully determined by the crossing-aymmetry requirement what in discussed in detail in $B$.
$F(\nu) \quad$ introduced in (19) is an analytical function of $\quad v$ with the cute along $\nu_{1}<\nu<\infty$ and $\rightarrow<\nu<-\nu$, and a (kinematic) pole at

$$
W^{2}=v^{2}+2 v v=0
$$

The crossing-symmetry requirements lead to

$$
F(-\nu)=\frac{y^{2}}{y^{2}+2 \mathbb{N}}-F^{2 N}(\nu) .
$$

and for amall

$$
F(\nu)=-\frac{a^{2}}{\nu}\left(1-\frac{2 \nu}{N}\right)+O\left(\nu^{2}\right)
$$

By applying the Cauchy formula with a contour dravin in Fig. 6 with $\rho \rightarrow \infty$, to $F\left(v_{0}\right)$ and writing down the dispersion relatione with subtraction, we get

$$
\begin{aligned}
& +\frac{\nu_{0}^{2}}{2 \pi i} \int_{e_{0}} F\left(\nu+\nu_{0}^{2}\left(\nu-v_{6}\right) d \nu+\frac{\nu_{0}^{2}}{2 \pi \int_{c_{-}}} \frac{F(\nu-i 6)}{\nu^{2}(\nu-\nu)} d \nu\right.
\end{aligned}
$$

and

$$
\begin{equation*}
A \rho F\left(v_{0}\right)=\frac{1^{2}}{N}\left(1-2 \nu_{0}\right), X\left(\nu_{0}\right) \cdot \frac{4 v_{0}^{2} R R_{e} F\left(v_{0}+N\right)}{N\left(\nu_{0}\right)} \tag{24}
\end{equation*}
$$

whare

$$
\begin{equation*}
R\left(\nu_{0}\right)=\frac{v_{0}^{\prime}}{\pi} \sum_{\nu_{1}} \frac{I m}{\nu^{2}}[\nu)\left[\frac{1}{\nu-\nu_{0}}+\frac{N^{2}+2 N \nu}{T^{2}-2 N \nu}-\frac{1}{\nu+\nu_{0}}\right]!I_{\nu} \tag{25}
\end{equation*}
$$

Since

$$
K(M / 2)=0,
$$

it is impossible to determine $\Pi \mathrm{me} \mathrm{F}(1 / 2)$ from (24), and this quantity enters as a free parameter which is necessary to determine from the experimental data. If we restrict ourselves to the photoproduction in the states with $J \leqslant 3 / 2$, then

$$
\begin{aligned}
& \operatorname{lm} F(\nu)=\nu \left\lvert\, \frac{M}{W}\left(\left|E_{1}\right|^{2}-\left|M_{2}\right|^{2}\right)+2\left(\left|E,\left.\right|^{2}-\left|M_{3}\right|^{2}\right)\left(1+1 / 2 \frac{W-W}{W}\right)+\right.\right. \\
& \left.\quad+1 H_{4} \frac{W+H}{W}\left(\left|E_{2}\right|^{2}-\left|M_{2}\right|^{2}\right)-\frac{W-M}{W} \operatorname{Re}\left(E_{3}^{*} H_{2}-E_{2} M_{3}^{*}\right) \right\rvert\,
\end{aligned}
$$

In Fig. 7 are given the results of the estimate of $\mathrm{Re}\left(\mathrm{R}_{1}-\mathrm{R}_{2}\right)$ (24), the contribution proportional to Re F (M/2) being neglected.

As is seen from (4. B), in order to estimate $R_{5}-\Pi_{6}$ at $Q^{2}=0$ it suffices to consider the function

the dispersion relations for which are of the form

$$
\begin{equation*}
R e b\left(\nu_{0}\right)-\psi(0)=\frac{3 \nu_{0}^{2}}{\pi} \int_{\nu_{t}}^{\infty} \frac{\operatorname{lm} \psi(\nu) d \nu}{\nu\left(\nu 2-\nu_{0}^{2}\right)}, \tag{27}
\end{equation*}
$$

where, as it follows from (2.B)

$$
\begin{equation*}
\psi(\eta)=-\frac{e^{2}}{M} \cdot \frac{2+\lambda}{2} \tag{28}
\end{equation*}
$$

and $\ln \psi(\nu)=\left(\frac{W}{M}\right)^{2}\left\{-W\left[\frac{l}{6}\left(\left|E_{2}\right|^{2}-\left|M_{2}\right|^{2}\right\}+\operatorname{Re}\left(E_{2}^{*} M_{3}-M_{2}^{*} E_{3}\right)\right]+\frac{U \nu}{(H+W)}\left[3\left(\left|E_{y}\right|^{2}-\left|H_{y}\right|^{2}\right)+\right.\right.$

$$
\begin{equation*}
\left.\left.+1 / 4\left(\left|E_{2}\right|^{2}-\left|M_{2}\right|^{2}\right)+\operatorname{Re}\left(E_{2}^{*} 甘_{5}-H_{2}^{*} E_{3}\right)\right]\right\} \tag{29}
\end{equation*}
$$

The results of the estimation for $\mathrm{Re}\left(\mathrm{R}_{5}-\mathrm{R}_{6}\right)$ at $\mathrm{Q}^{2}=0$ are given in Fig. 4 . The estimates of the quantities $\mathrm{R}_{3}+\mathrm{R}_{4}$ and $\mathrm{n}_{5}-\mathrm{R}_{6}$ which play the main part in the differential cross section at $\nu \geqslant 1$ are not much different from those obtained earlier ${ }^{/ 6 /}$.

These results are of interest for studying the dependence of the amplitudes near the threshold for a new reaction. ${ }^{/ 6 /}$. In the case under consideration all the estimates can be made completely. We stress the dependence of the amplitude $\operatorname{Re}\left(\Pi_{1}+R_{2}\right)$ the value of which keeps on decreasing higher than the thres-
hold. This result prints out that for other processes the sharp dependence of the imaginary parts of the amplitudes higher than the threshold may lead to the displacement of the threshold minimum (or maximam) of the cross section from the reaction threshold.

The results of the calculations with the aid of $A_{1}(\nu, 0)$ for the angular distributions

$$
\sigma(\theta)=\sum_{t=0}^{3} \quad a_{1} \cos ^{1} \theta
$$

for $\theta=90^{\circ}, 135^{\circ}, 139^{\circ}$ and $180^{\circ}$, as well as for the total elastic scattering crose section

$$
\frac{\sigma_{*}}{4 \pi}=D_{0}+1 / 2 D_{2}
$$

and for the recoil nucleon polarization at $\theta=90^{\circ}$ are presented in Figs. 5, 8-12.
The experimental data are summarized in $/ 10 /$ and $/ 17 /$. The coefficient

$$
B_{g}\left(\nu_{0}\right)=2\left[\left|R_{g}+R_{6}\right|^{2}-\left|R_{g}-R_{g}\right|^{2}\right]
$$

is close to zero throughout the energy region $v_{0} \leqslant 2$.
The experimental data are seems to indicate, that the quantity $\operatorname{Re}\left(\mathrm{R}_{5}-\mathrm{R}_{6}\right)$ is positive. However, we could not do it by introducing Re $\mathrm{F}(\mathrm{M} / 2) \neq 0$. The requirement for $\mathrm{Re}\left(\mathrm{R}_{5}-\mathrm{R}_{6}\right)$ to be positive leads to large (negative) values of $\operatorname{ReF}(\mathrm{M} / 2)$ what increases considerably the contribution
$R_{1}-\left.R_{2}\right|^{2}$ to the cross aection and fails to improve the agreement with the experimental data.
It is worth while noting that outside the region $1<\nu_{0}<1.3$ there is a satisfactory agreement between the diapersion analyais and experimental data. In the region $1<\nu_{0}<1.3$ especially sensitive to the dispersion effects if is, probably, necessary to consider the contribution of higher states for which more information on the pion photoproduction in the wide energe region is required.

| $V_{0}$ | 0 | 0,5 | 0,8 | 0,90 | 0,95 | 0,97 | I,00 | I, 05 | I, IO | I,20 | I,50 | 2,00 | 2,50 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $\operatorname{Re}\left(R_{s}+R_{2}\right)$ | -I | -0,8 | -0,55 | -0,38 | -0,20 | -0,23 | -0,17 | -0,14 | -0,12 | -0,09 | $-0,06$ | -0,16 | -0,52 |
| $\operatorname{Im}\left(R_{1}+R_{2}\right)$ | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0,I5 | 0,23 | 0,35 | 0,83 | 2,83 | 2,21 |
| $\left\|R_{1}+R_{2}\right\|^{2}$ | I | 0,64 | 0,30 | 0,14 | 0,08 | 0,05 | 0,03 | 0,04 | 0,06 | 0,13 | 0,69 | 8,0 | 5,15 |
| $\operatorname{Re}\left(R_{1}-R_{2}\right)$ | -I | -0,74 | -0,54 | -0,44 | -0,38 | -0,34 | -0,20 | -0,23 | -0,26 | -0,29 | -0,43 | -0,05 | +1,71 |
| $\operatorname{Im}\left(R_{1}-R_{2}\right)$ | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0,14 | 0,20 | 0,26 | 0,22 | -I,03 | -0,48 |
| $\left\|R_{1}-R_{2}\right\|^{2}$ | I | 0,54 | 0,29 | 0,19 | 0,15 | 0,12 | 0,04 | 0,07 | 0,II | 0,15 | 0,23 | I,06 | 3,15 |
| Re $\left(R_{3}+R_{4}\right)$ | 0 | -0,29 | -0,43 | -0,44 | -0,43 | -0,41 | -0,34 | -0,42 | -0,50 | -0,68 | -I,I2 | -I,78 | -I,70 |
| $\operatorname{Im}\left(R_{3}+R_{4}\right)$ | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0,14 | 0,21 | 0,28 | 0,38 | -0,15 | +0,13 |
| $\left\|R_{3}+R_{4}\right\|^{2}$ | 0 | 0,08 | 0,18 | 0,19 | 0,18 | 0,17 | 0,I2 | 0,20 | 0,25 | 0,54 | I,4I | 3,20 | 2,87 |
| $\operatorname{Re}\left(R_{3}-R_{4}\right)$ | 0 | 0,26 | 0,50 | 0,60 | 0,68 | 0,72 | 0,83 | 0,88 | 0,93 | I, OI | I, 12 | I,20 | 0,85 |
| $\operatorname{Im}\left(R_{3}-R_{4}\right)$ | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0,I5 | 0,27 | 0,32 | 0,63 | I,6I | I, 35 |
| $\left\|R_{3}-R_{4}\right\|^{2}$ | 0 | 0,07 | 0,25 | 0,36 | 0,46 | 0,52 | 0,69 | 0,80 | 0,94 | I, I2 | I,65 | 4,04 | 2,54 |
| $R_{e}\left(R_{5}+R_{6}\right)$ | 0 | 0,08 | 0,II | 0,II | 0,II | 0,II | 0,12 | 0,12 | 0,10 | 0,II | 0,13 | 0,25 | -0,7 |
| $\left\|R_{5}+R_{6}\right\|^{2}$ | 0 | 0,006 | 0,01 | 0,OI | 0,01 | 0,0I | 0,01 | 0,01 | 0,OI | 0,OI | 0,OI | 0,06 | 0,49 |
| Re $\left(R_{5}-R_{6}\right)$ | 0 | -0,08 | -0,10 | -0,II | -0,II | -0,II | -0,12 | -0,11 | -0,II | -0,1I | -0,II | -0,09 | -0,08 |
| $\left\|R_{5}-R_{6}\right\|^{2}$ | 0 | 0,006 | 0,01 | 0,01 | 0,01 | 0,01 | 0,0I | 0,01 | 0,01 | 0,OI | 0,OI | 0,01 | 0,006 |
| $\operatorname{Re} \overline{\left(R_{3}+R_{4}+2 \vec{R}_{5}+2 R_{6}\right)}$ | 0 | -0,13 | -0,20 | -0,21 | -0,20 | -0,18 | -0,10 | -0,19 | -0,26 | -0,45 | -0,86 | -I,29 | -3,06 |
| $\left\|R_{3}+R_{4}+2 R_{5}+2 R_{6}\right\|^{2}$ | 0 | 0,02 | 0,04 | 0,05 | 0,04 | 0,03 | 0,01 | 0,06 | 0,II | 0,28 | 0,90. | 1,68 | 9,36 |



Fig. 1


Fig. 2


Fis. 3


Fig. 7


Fis. 5


Fig. 6


Fig. 7


Fig. 8


Fig. 9


Fis. 10


Fig. 11


Fis. 1 ?

## neferences

1. M. Gell-Mann, M.L. Goldberger, W.E. Thirring. Phys.Rev., 95, 1612, 1954.
2. H Н. Боголюбов, ㅈ.B. Ширков. ДАН СССР, 113, 529, 1057. A. А. Логунов. Диссертаиия $195 \theta$ год.
3. M. Cini, R. Stroffolini, Nucl.Phys., 5, 684, 1958.
4. R.H. Capps. Phys. Rev., 106. 31, 1957; 108, 1032, 1957.
5. T. Akiba, J. Sato. Prog. Theor. Phys., 19, 93, 1958.
6. Л, Н. Лапидус, Чжоу Гуан-чжао. жЭТФ, 37, 1714, 1959; 38, 201, 1960 .
7. V. Jacob, J. Mathews. Phys.Rev., 117, 854, 1960.
8. V. Gell-Mann, M.I. Goldherger. Proc. 1954 Glasgow Conf. on Nucl, and Meson Physics. Pergamon Press London Y.Y.
9. L.G. Hyman, R. Ely, D.H. Frish, M.A. Wahlig. Phys.Rev. Lett, 3, 93, 1959.
10. G. Bernardini, A.O. Yanson, A.C. Odian, T. Yamagata, L. A. Auerbach, I. Filosofo. Nuoco Cimento, 18, 1203 , 1959.
11. F. E. L.ow. Proc. 1958 Ann. Internat. Conf. on Energy Physics. P. 98.
12. G. Glasser, N. Seeman, B. Stiller, Bull. Am. Phys.Soc. 6, 1, 1961.
13. П.И. Лапидус, Чжоу Гуан-чжао. Preprint ОИЯИ Д-681, ЖЭТФ /in press/. Further is reffered as peper "A"
14. Л.М. Лапидус, Чжоу Гуан-чжяо. Preprint ОНЯН, Д-682, ЖЭТФ / in press /. Further is reffered as paper ${ }^{*} B^{* *}$.
15. Y. Yamaguchi, Prog. Theor. Phys. 12, III, 1954. S. Vinardi, Y, Yamaguchi. Prog. Theor. Phys. 17, 651, 1957.
16. S. Winami. 'Photon-Proton Collision at $250-800 \mathrm{MeV}$ '. (Pr eprint).
17. Б.Б. Говорков, В.Н. Гольданскии, О.А.Кардухин, А.В. Куценко, В.В. В.В. Павловская. ЖЭТФ 38, 1695, $1860 . г$.
В.В. Павловская. Диссертация, 1 Ө61.

$$
\begin{gathered}
\text { Работа поступила в иядательскни отдел } \\
6 \text { мая } 1961 \mathrm{r} .
\end{gathered}
$$


[^0]:    * Thie result hes been obtalned also by W.N. Gribov.

[^1]:    * The mfgn meansthe differentiation with respeot to $Q^{2}$ and abbaequent teanatien to $Q^{2}=0$.

